

Sensor Characterisation with Darks

Merlin Fisher–Levine

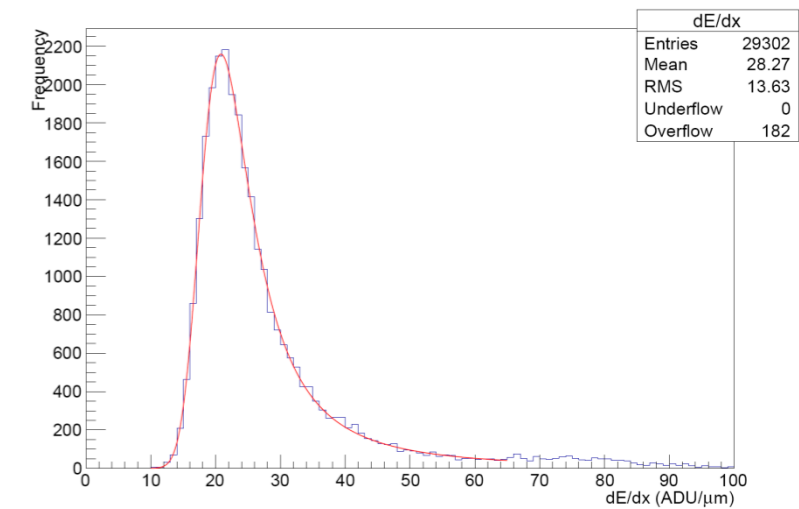
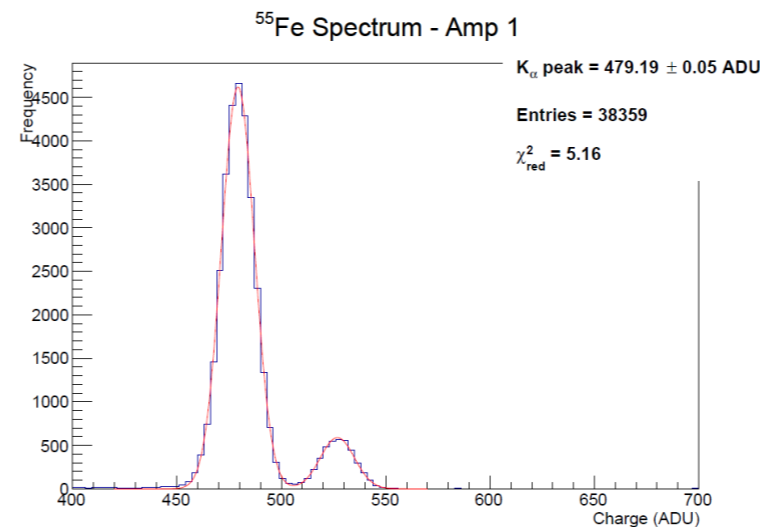
September 2014

Sensor characterisation

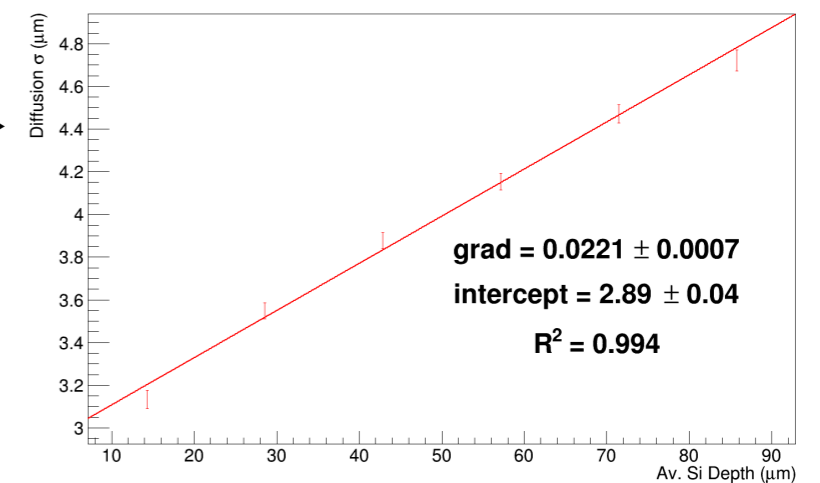
- ▶ Using dark exposures to characterise LSST sensors

- ▶ Measuring:

- ▶ Gain



- ▶ Diffusion



- ▶ Edge effects (track bending)

- ▶ Have been using LSST's DMStack for processing images

- ▶ With further analysis in python and ROOT (sorry)



Image assembly

- Image assembled from 16 amplifier sections
- Different ADC offsets cause large background ‘brightness’ variation
- Subtract these offsets to produce flat image

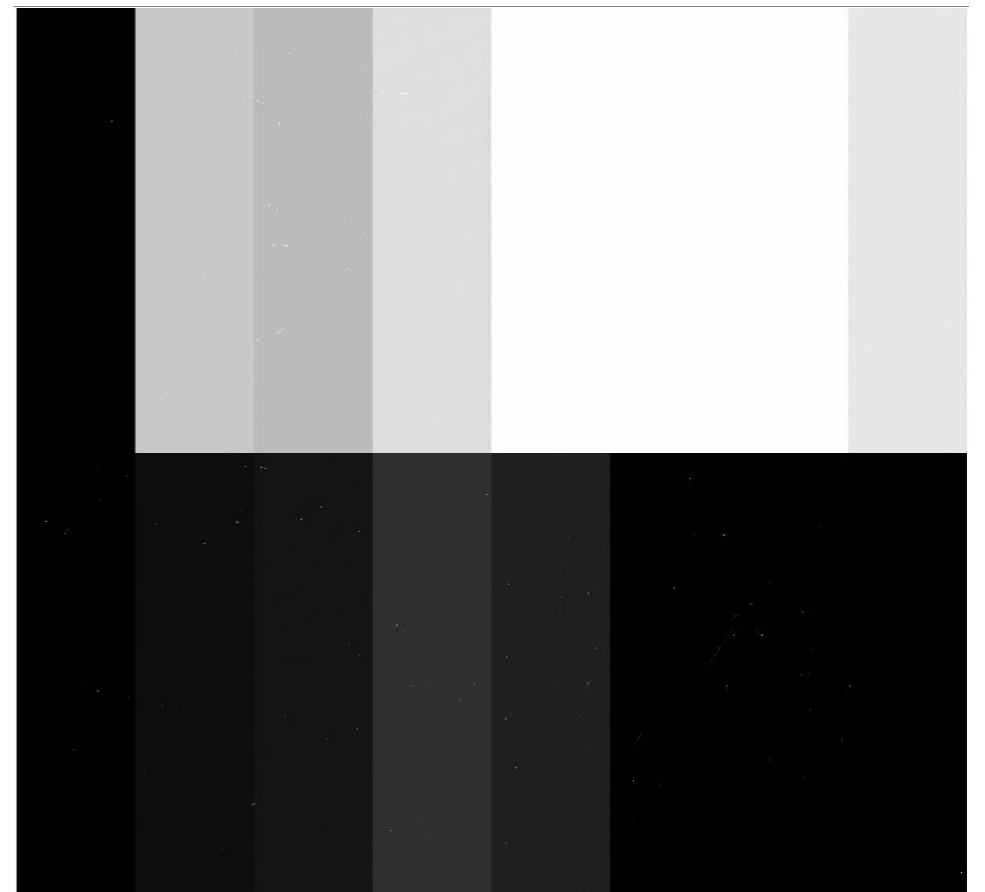


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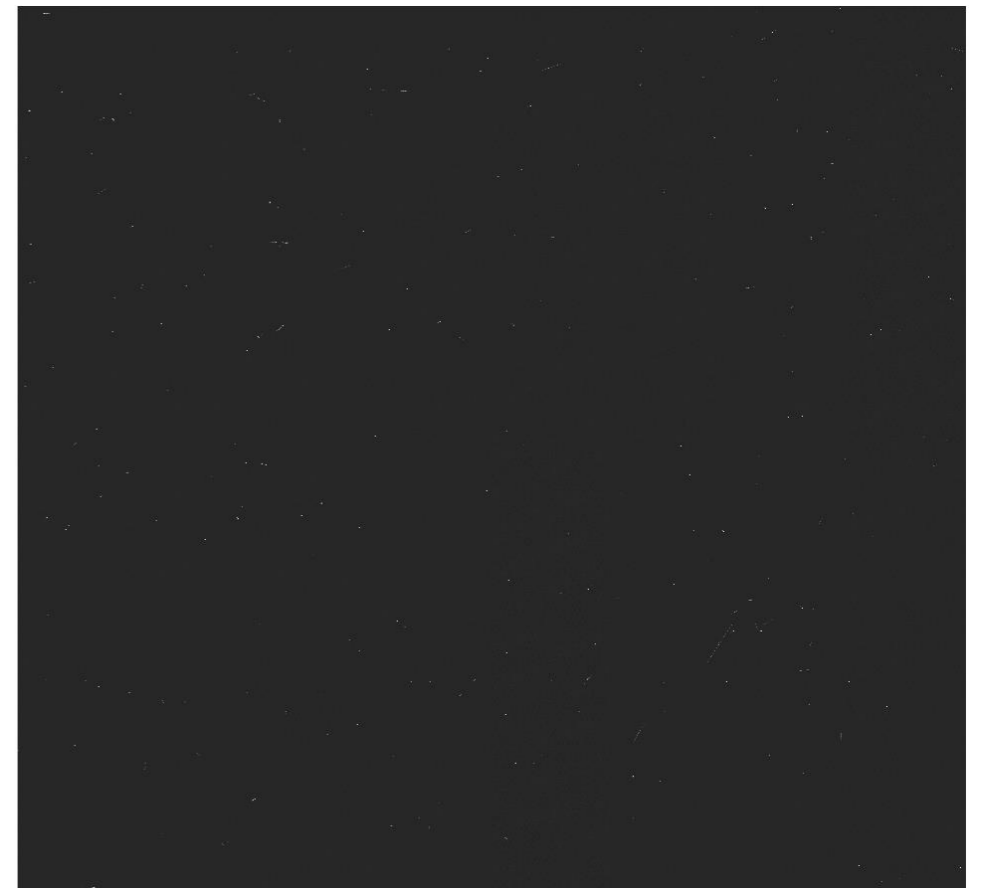


Image assembly

- Image assembled from 16 amplifier sections
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- Subtract these offsets to produce flat image
- Now use DMStack to find tracks in dark exposure

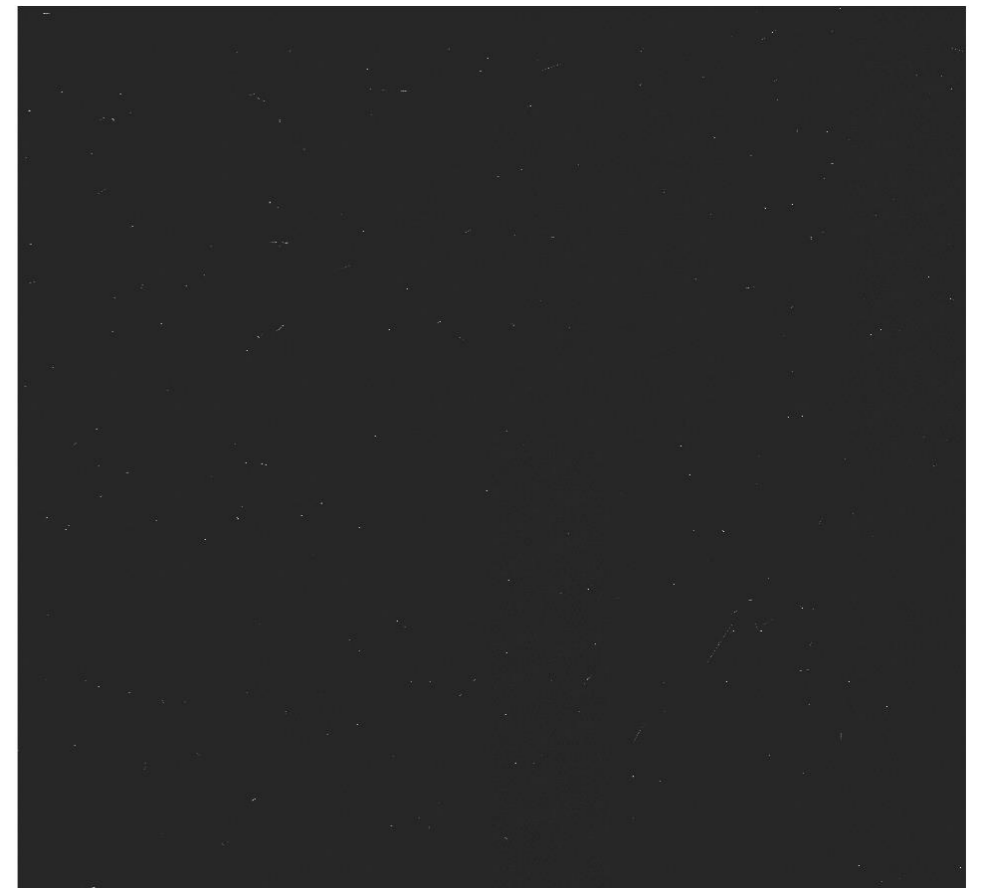
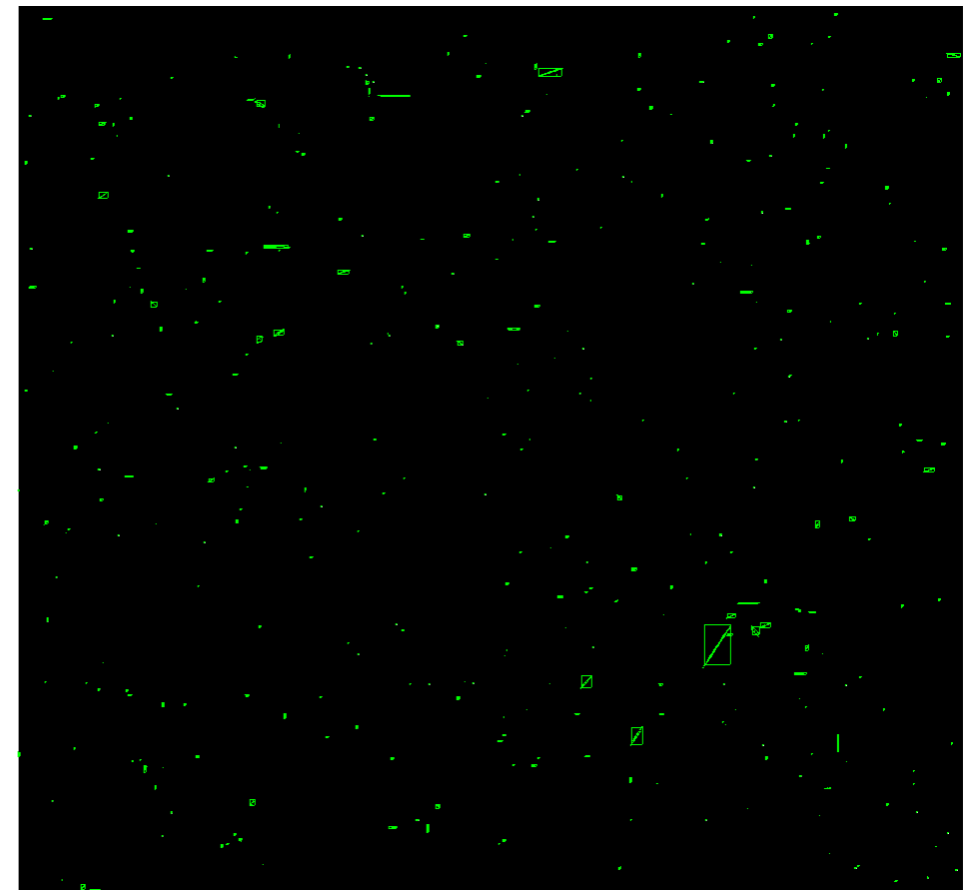


Image assembly

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- Different ADC offsets cause large background ‘brightness’ variation
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Track analysis for gain measurement

- Each tracks falls into 1 of 3 categories:
 - Worms
 - Spots
 - Cosmic ray muons

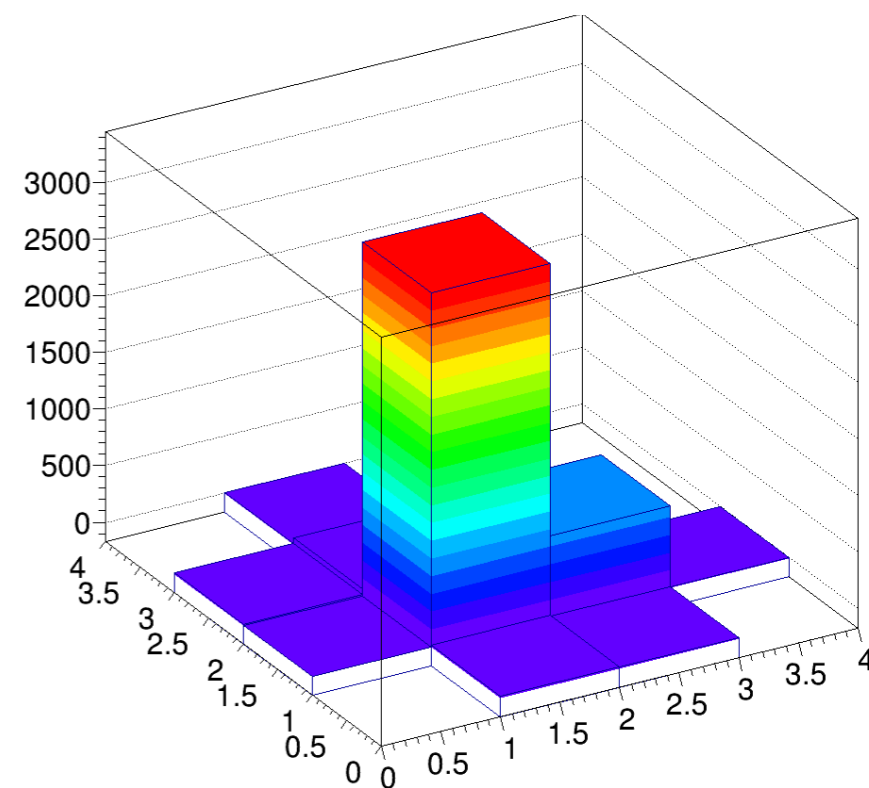
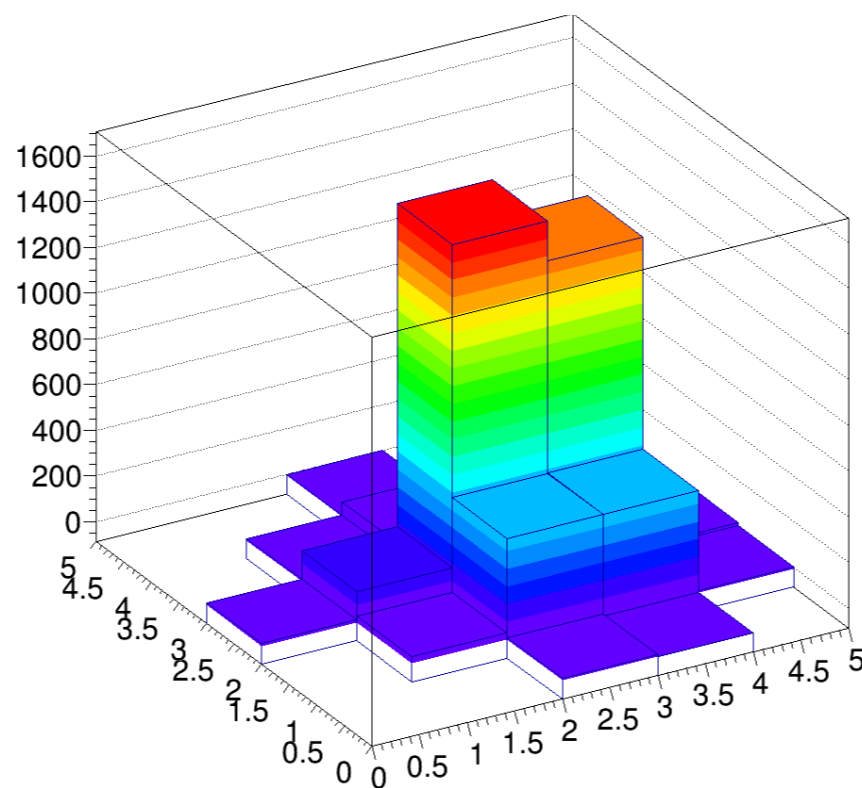
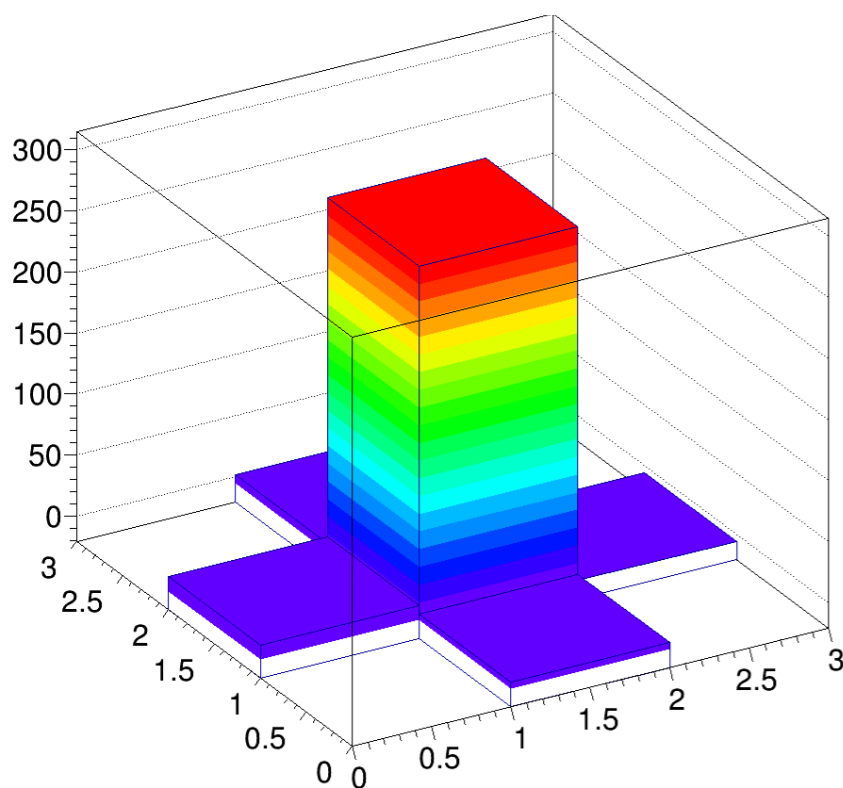
} Soft gammas – not useful – discard

← Useful, keep!
- Cosmics ray muons are “minimum ionising particles”
 - Energy per unit length is “constant” and well defined
 - Stochastic mechanism but has a well known distribution
 - Landau distributed energy loss
- Want to reject worms and spots, keeping cosmics...

Track cleaning

- Find all tracks in dataset
 - Cut dataset to only include long & straight tracks

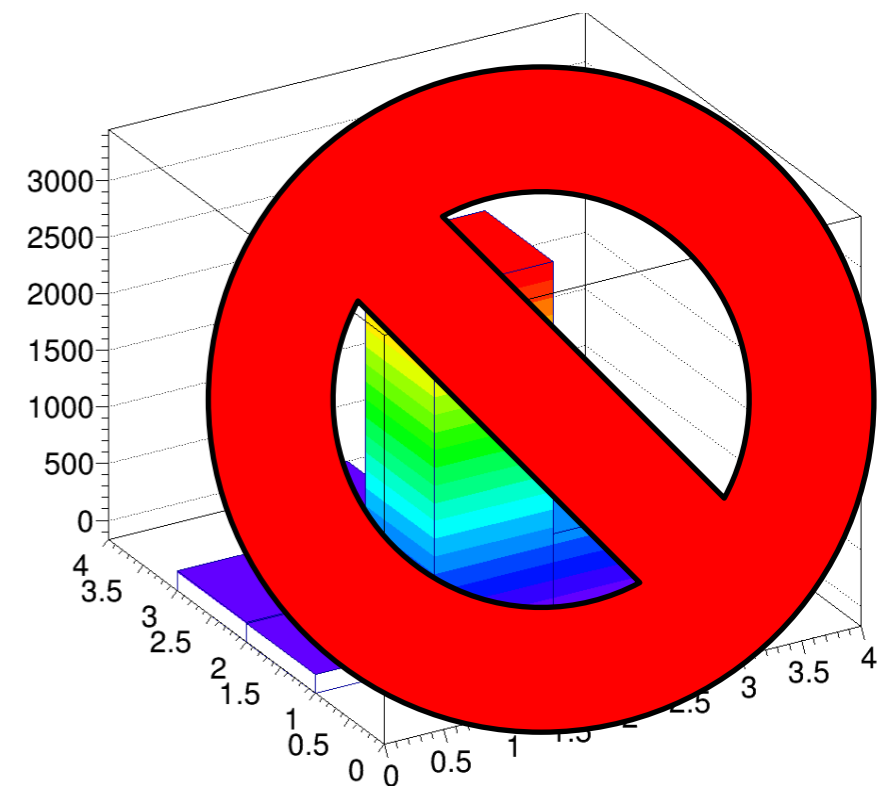
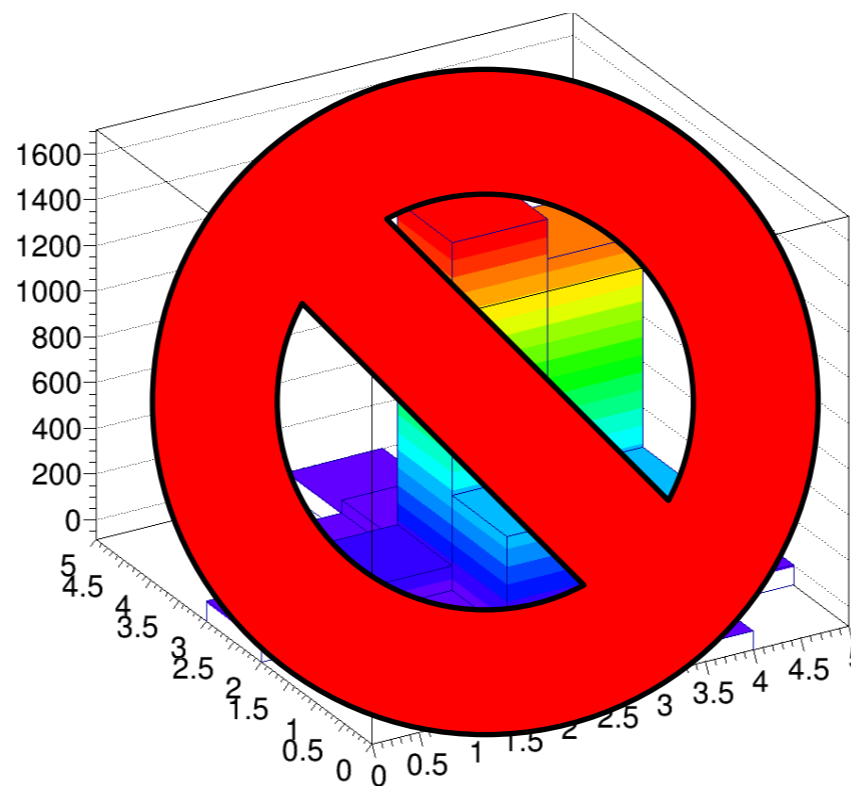
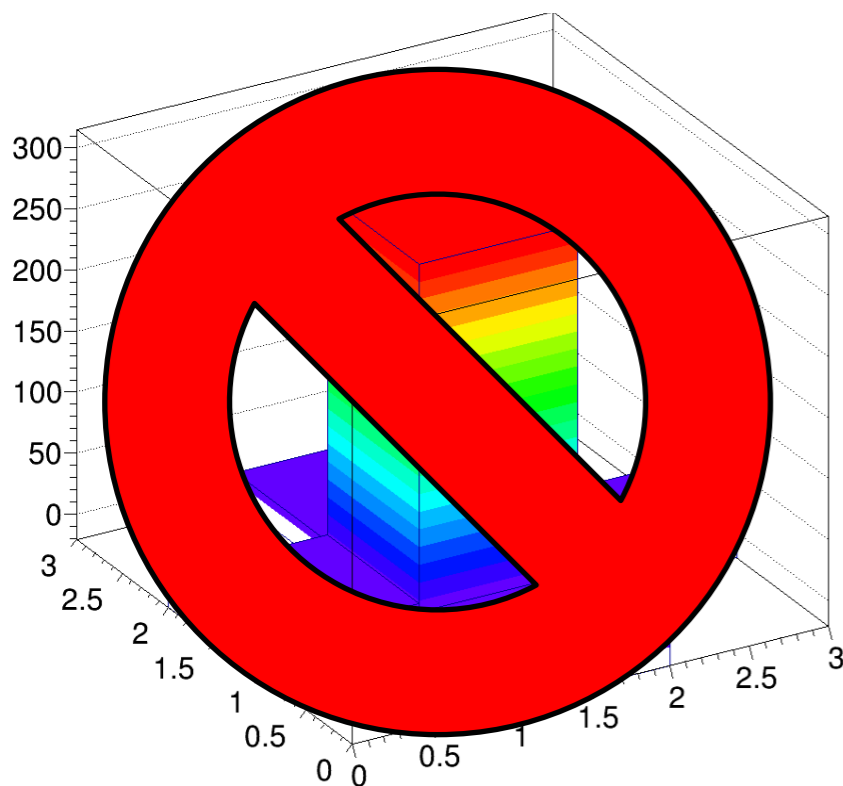
Remove spots



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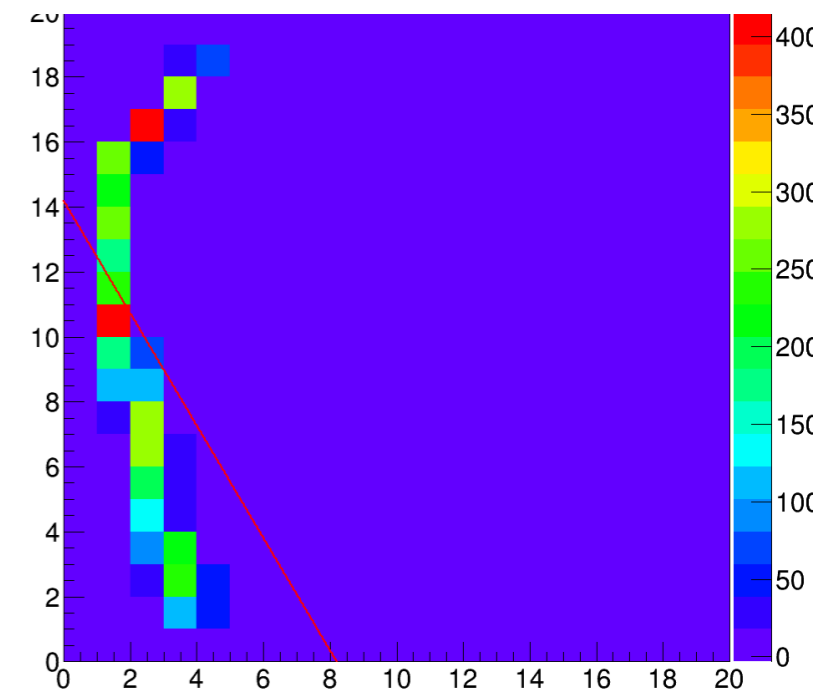
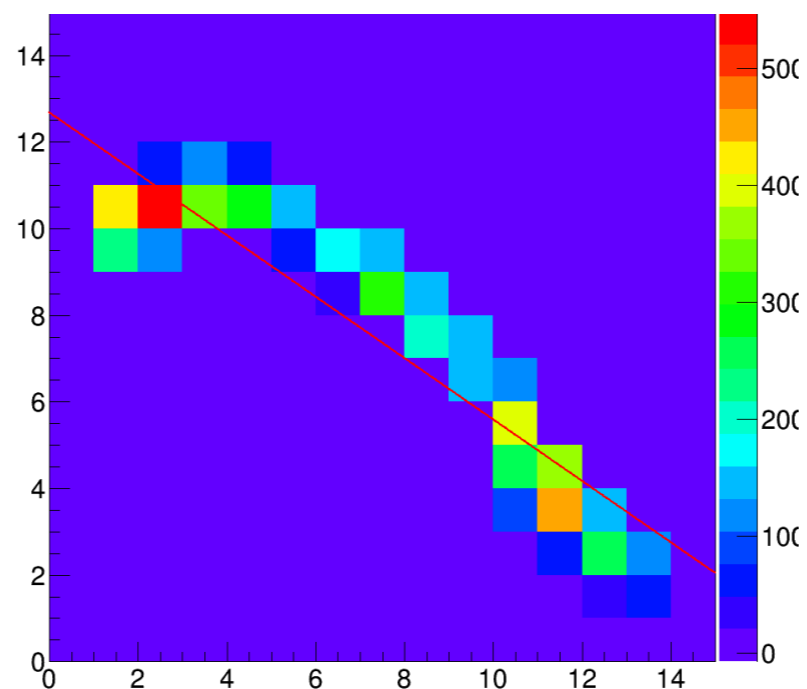
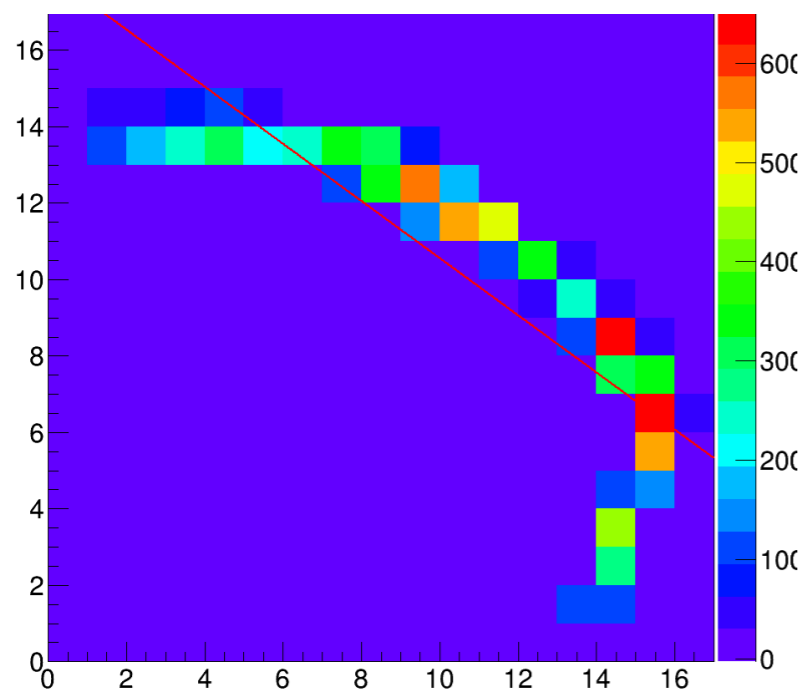
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Track cleaning

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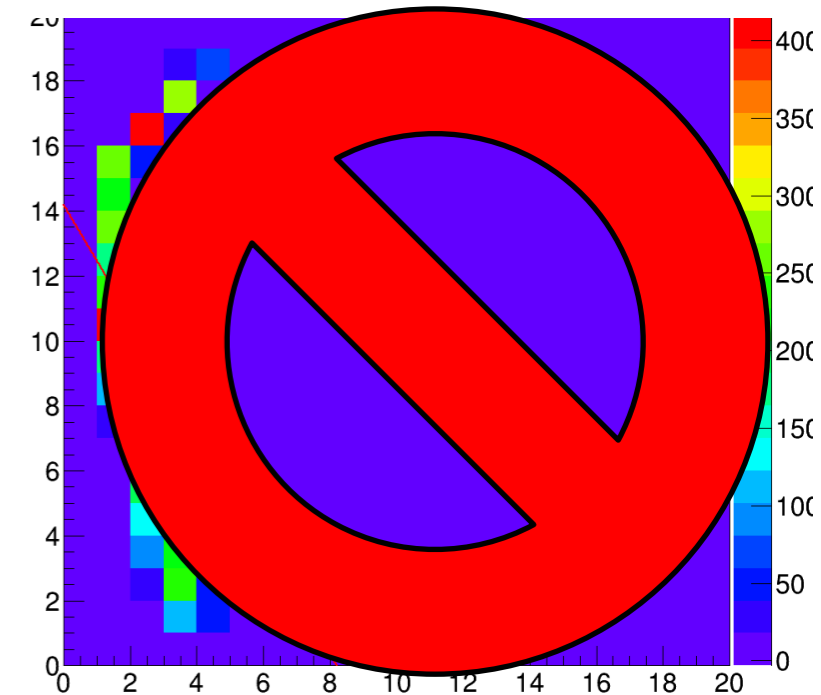
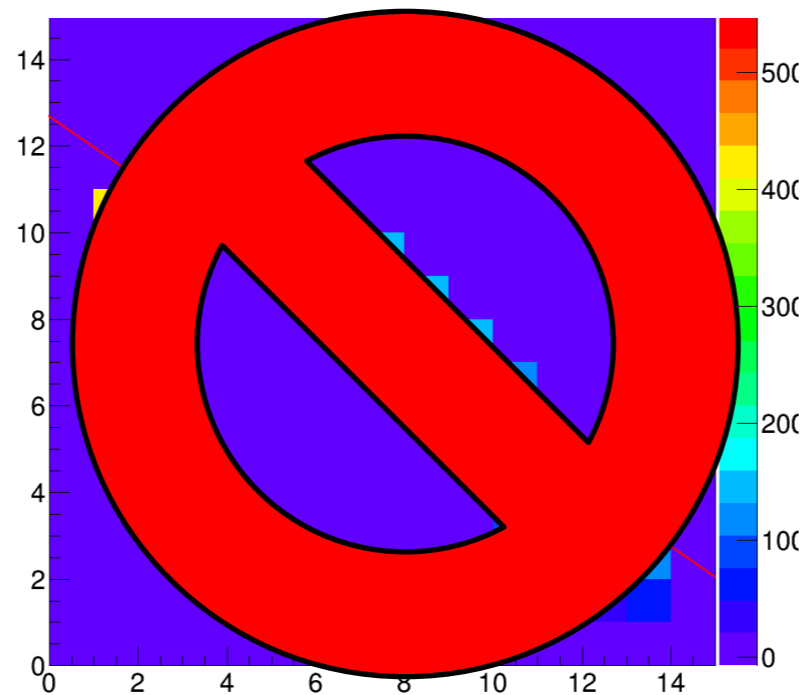
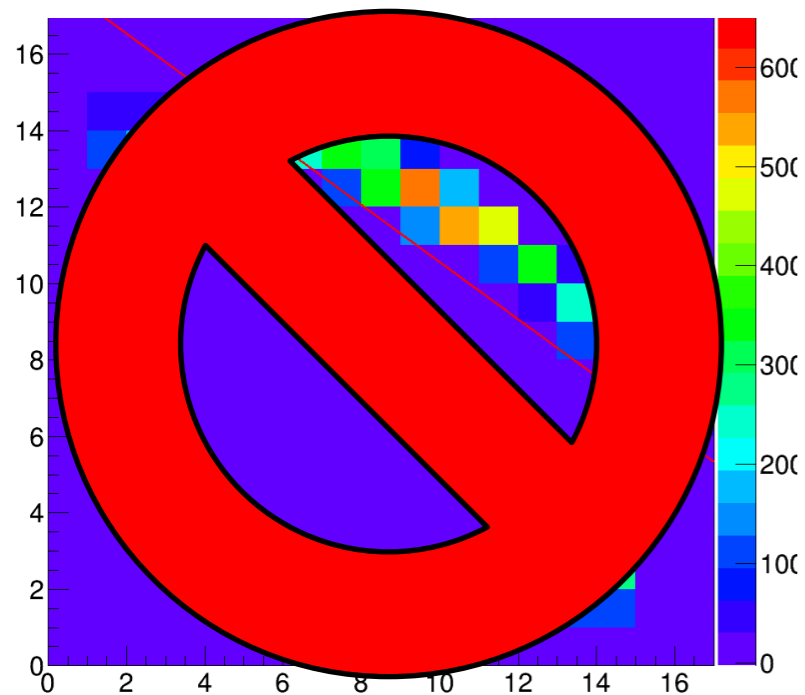
Remove worms



Track cleaning

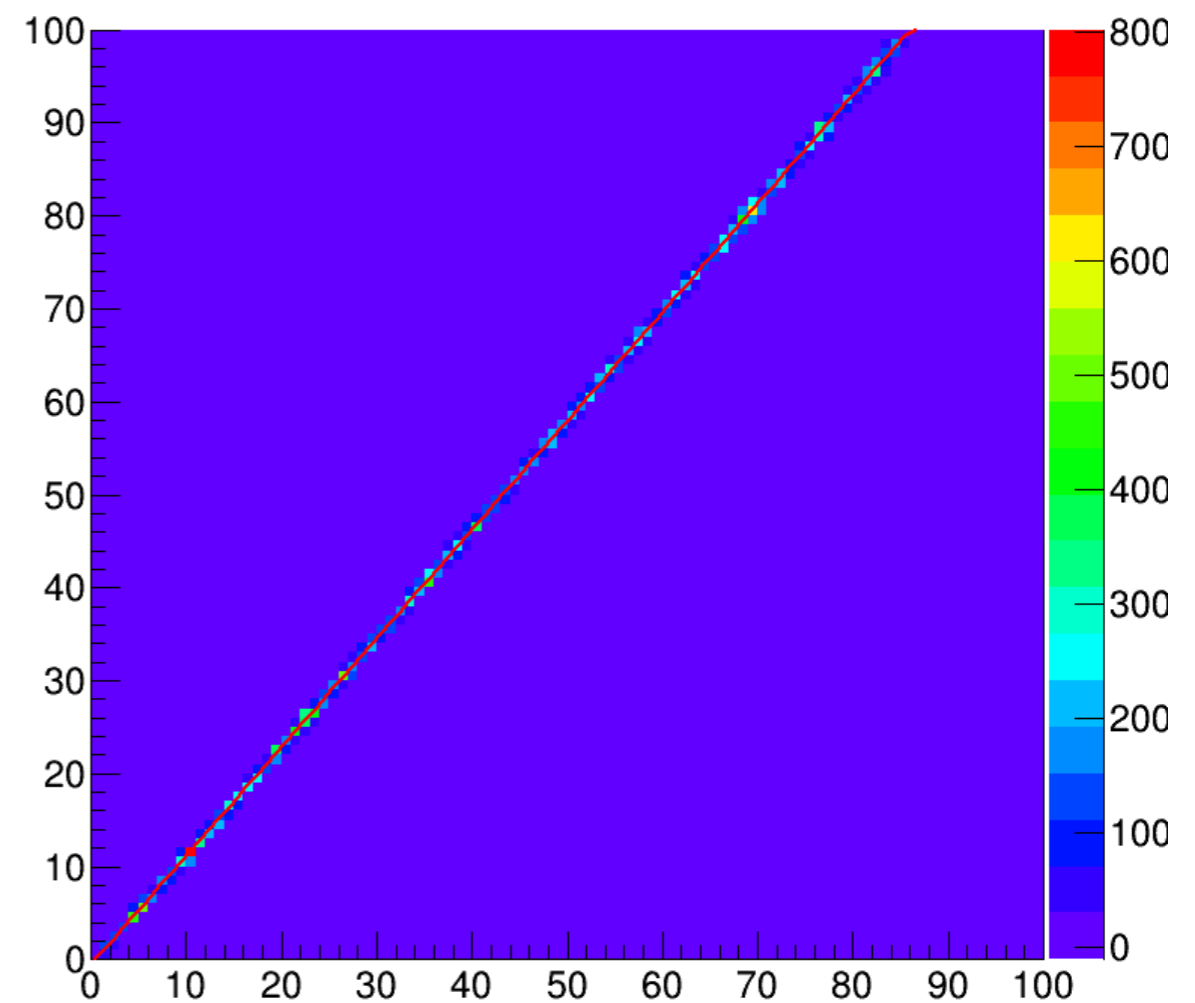
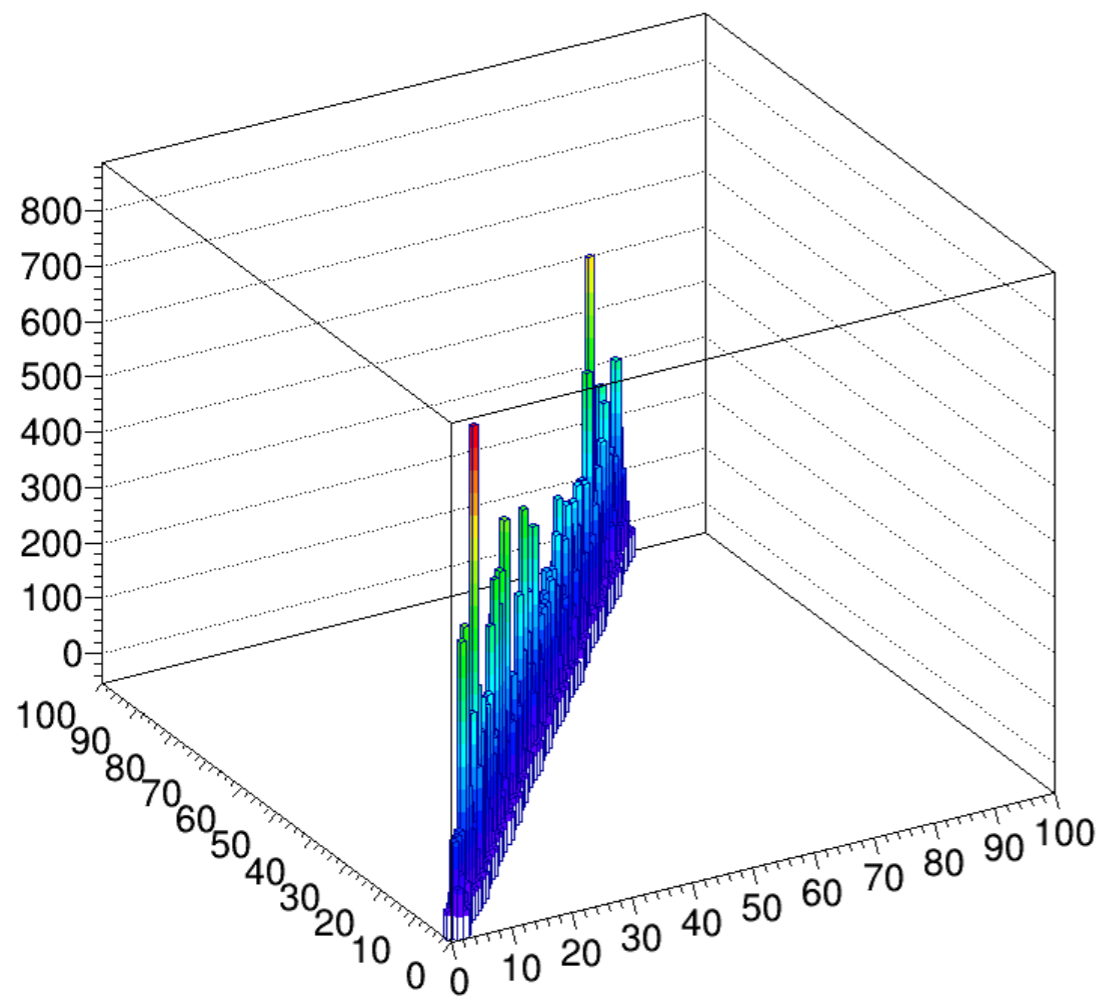
- Find all tracks in dataset
 - Cut dataset to only include long & straight tracks

Remove worms



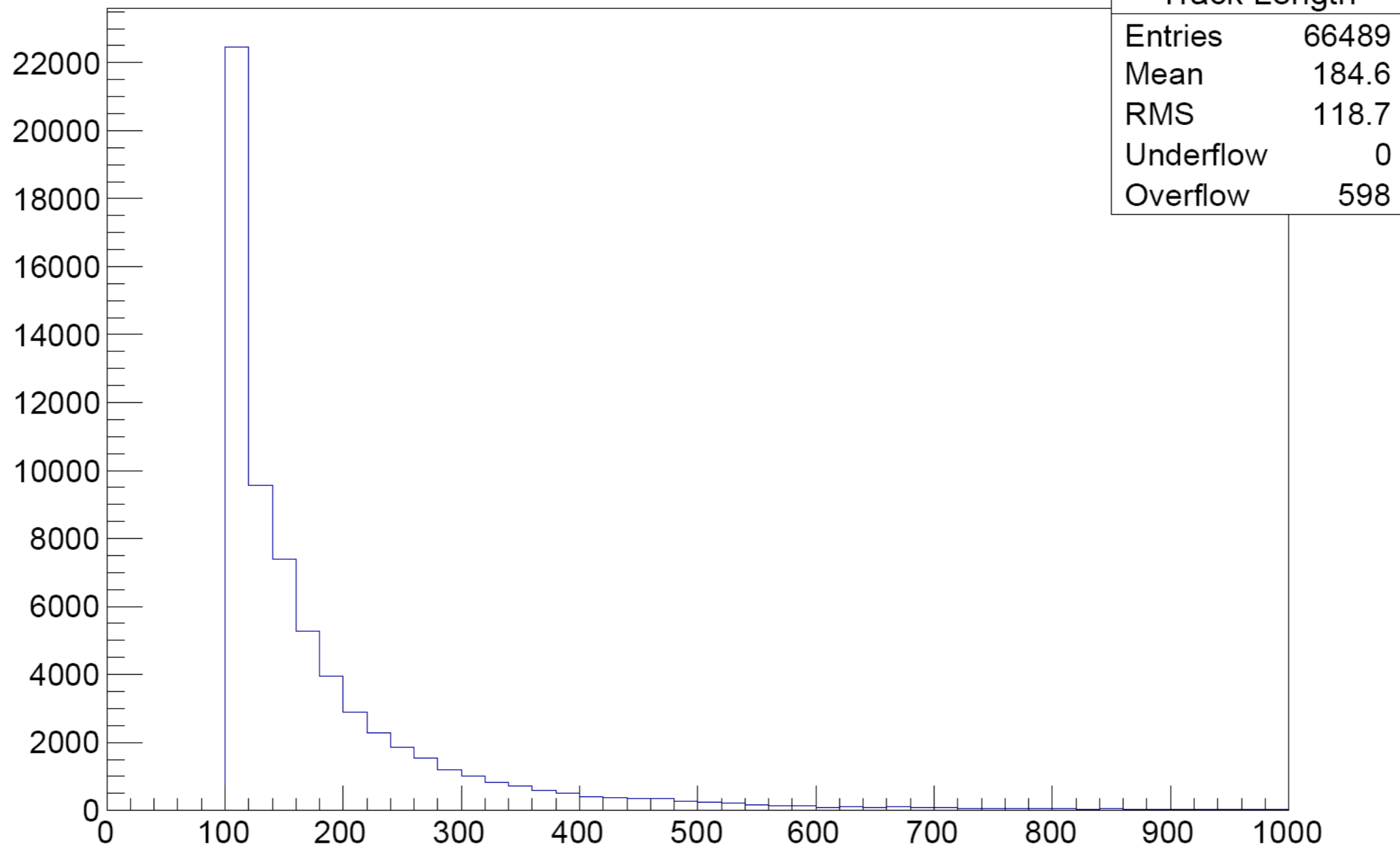
Track cleaning

- Dataset now only contains long straight tracks:



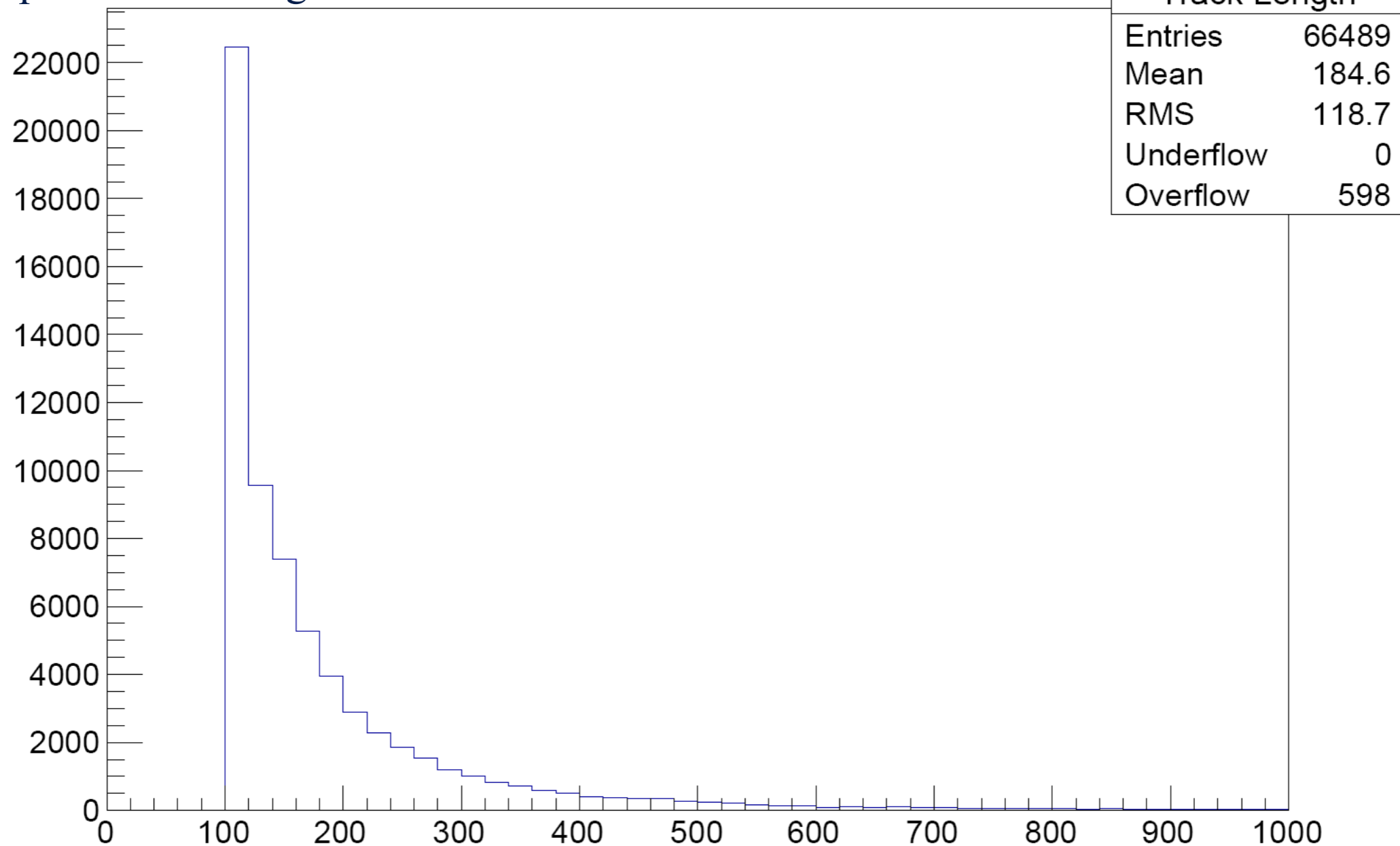
Track cleaning

- Track length distribution:
 - Defined by sensor thickness & orientation, combined with the angular distribution of cosmic ray muons
 - V. short tracks are due to spots and worms



Track cleaning

- Rejecting worms and spots, sum the total flux in each track
 - Histogram
- Produces a Landau distribution
 - Fit peak to extract gain...



Landau Distribution

Landau Distribution

$$f(x, \Delta) = \phi(\lambda) / \xi$$

$$\phi(\lambda) = \frac{1}{\pi} \int_0^\infty e^{-u \ln u - u \lambda} \sin(\pi u) du$$

$$\lambda = \frac{1}{\xi} [\Delta - \xi (\ln \xi - \ln \epsilon + 1 - C)]$$

$$\ln \epsilon = \ln \frac{(1 - \beta^2) I^2}{2 m c^2 \beta^2} + \beta^2$$

$$K = 2\pi N_0 r_e^2 m_e c^2 = 0.1535 \text{ MeV cm}^2/\text{g};$$

x is the path length in g/cm²;

$r_e = \frac{e^2}{4\pi\epsilon_0 m_e c^2} = 2.8179 \times 10^{-13}$ cm is the classical electron radius;

m_e is the mass of the electron;

N_0 is Avagadro's number, 6.022×10^{23} ;

I is the mean excitation energy averaged over all electrons in eV;

Z is the atomic number of the medium;

A is the atomic weight of the medium;

ρ is the density of the medium;

z is the charge of the incoming particle;

$\beta = v/c$ is the ratio of the incoming particle to the speed of light;

γ is the Lorentz factor $\frac{1}{\sqrt{1-\beta^2}}$;

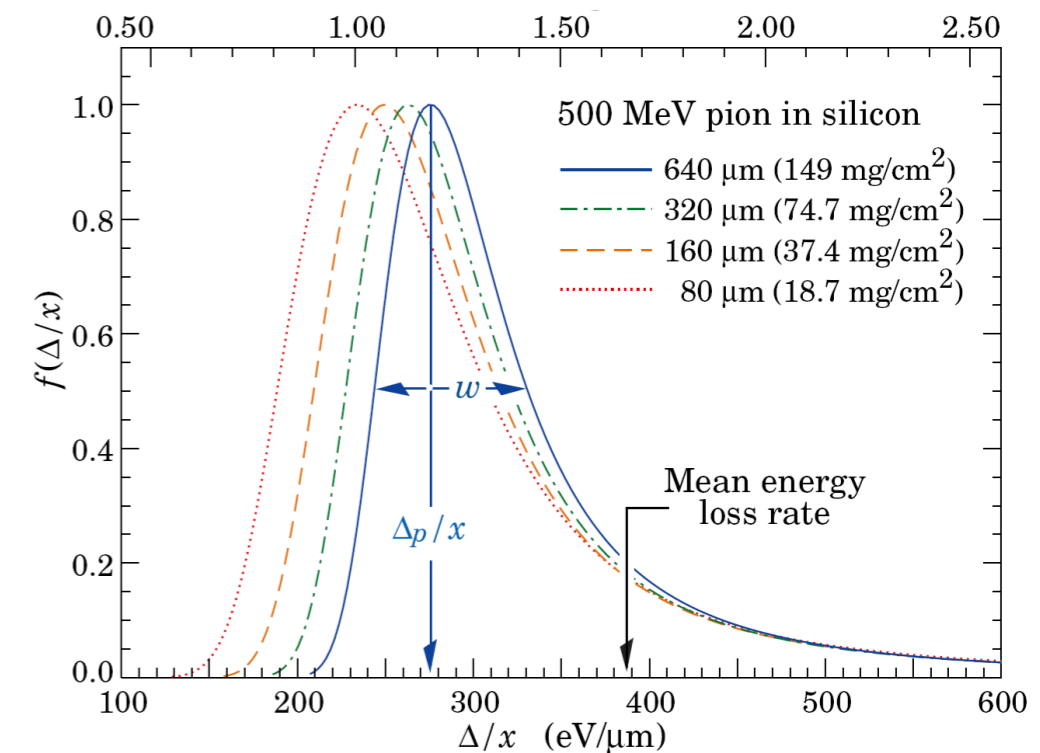
δ_{ion} is a density correction;

C is a shell correction;

Peak Energy loss

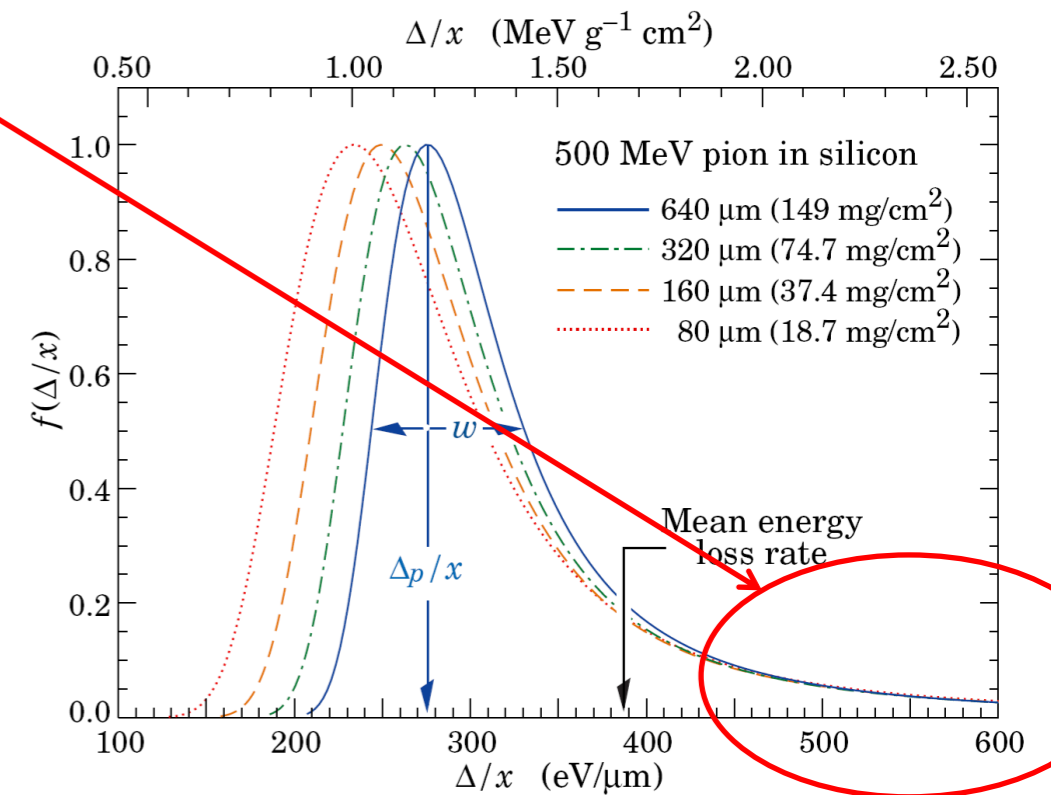
$$\Delta_p = \xi \left[\ln \frac{2 m c^2 \beta^2 \gamma^2}{I} + \ln \frac{\xi}{I} + j - \beta^2 - \delta_{ion}(\beta \gamma) \right]$$

$$\xi = x \cdot \frac{K}{2} \frac{Z}{A} \frac{1}{\beta^2} \text{ MeV}$$

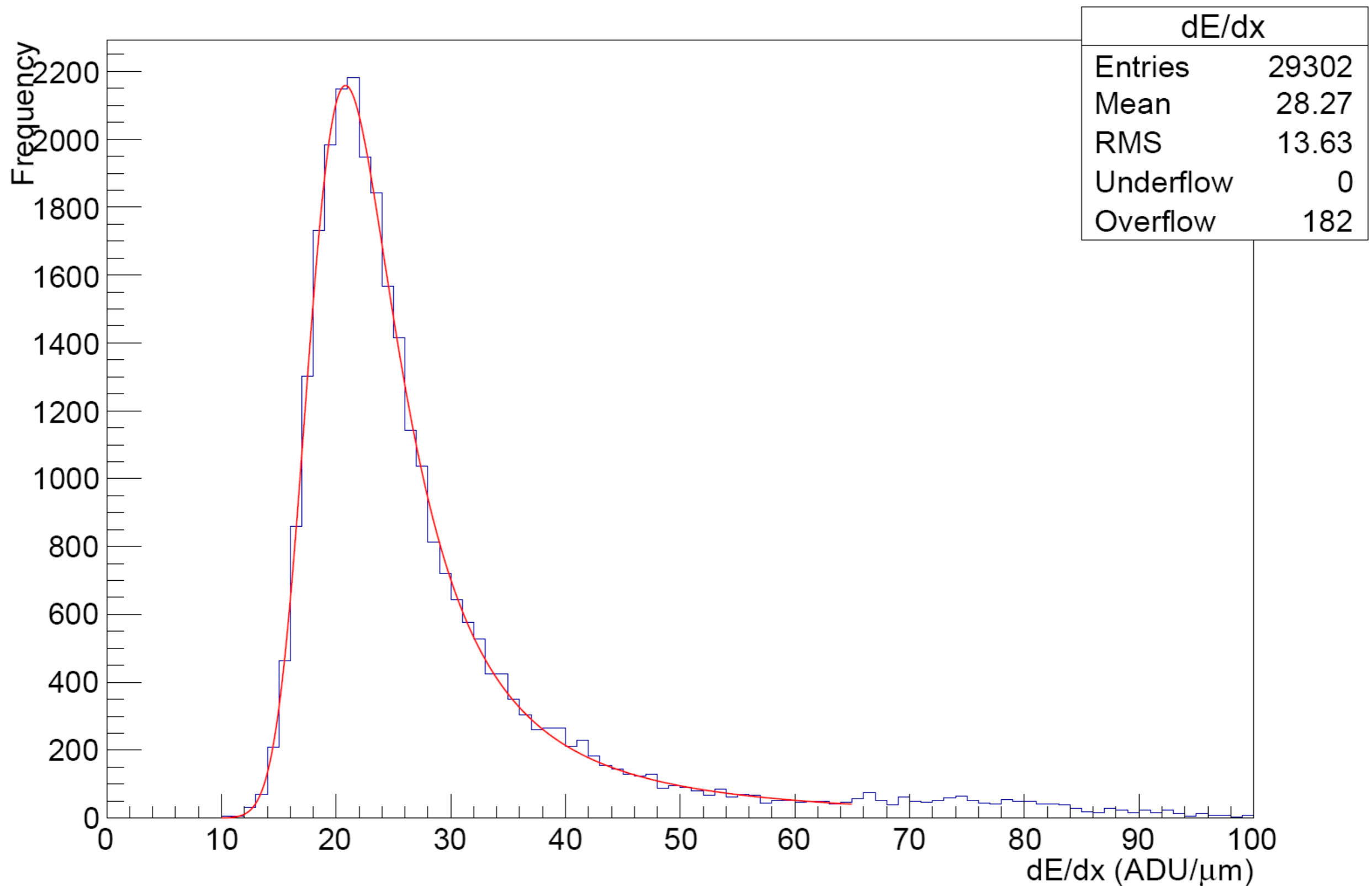


Landau Distribution

- ▶ Tail goes to “infinity”
 - ▶ (i.e. the energy of muon – several GeV!)
- ▶ This produces large δ -rays



Measured cosmic ray distribution



Muon energy spectrum - Amp 1

MPV = 18.81 ± 0.12 ADU

Entries = 1910.0

$\chi^2_{red} = 1.87$

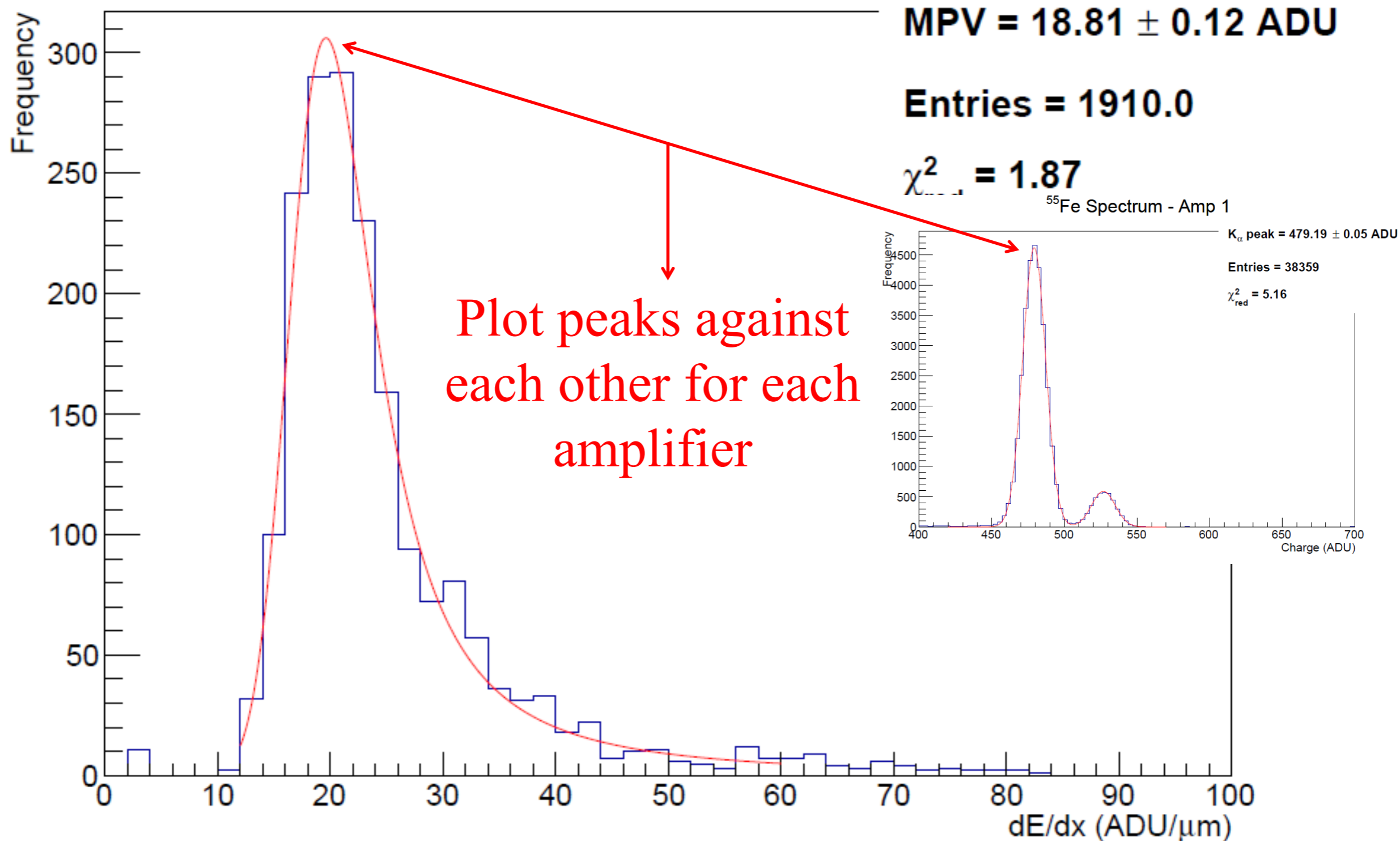
⁵⁵Fe Spectrum - Amp 1

K_α peak = 479.19 ± 0.05 ADU

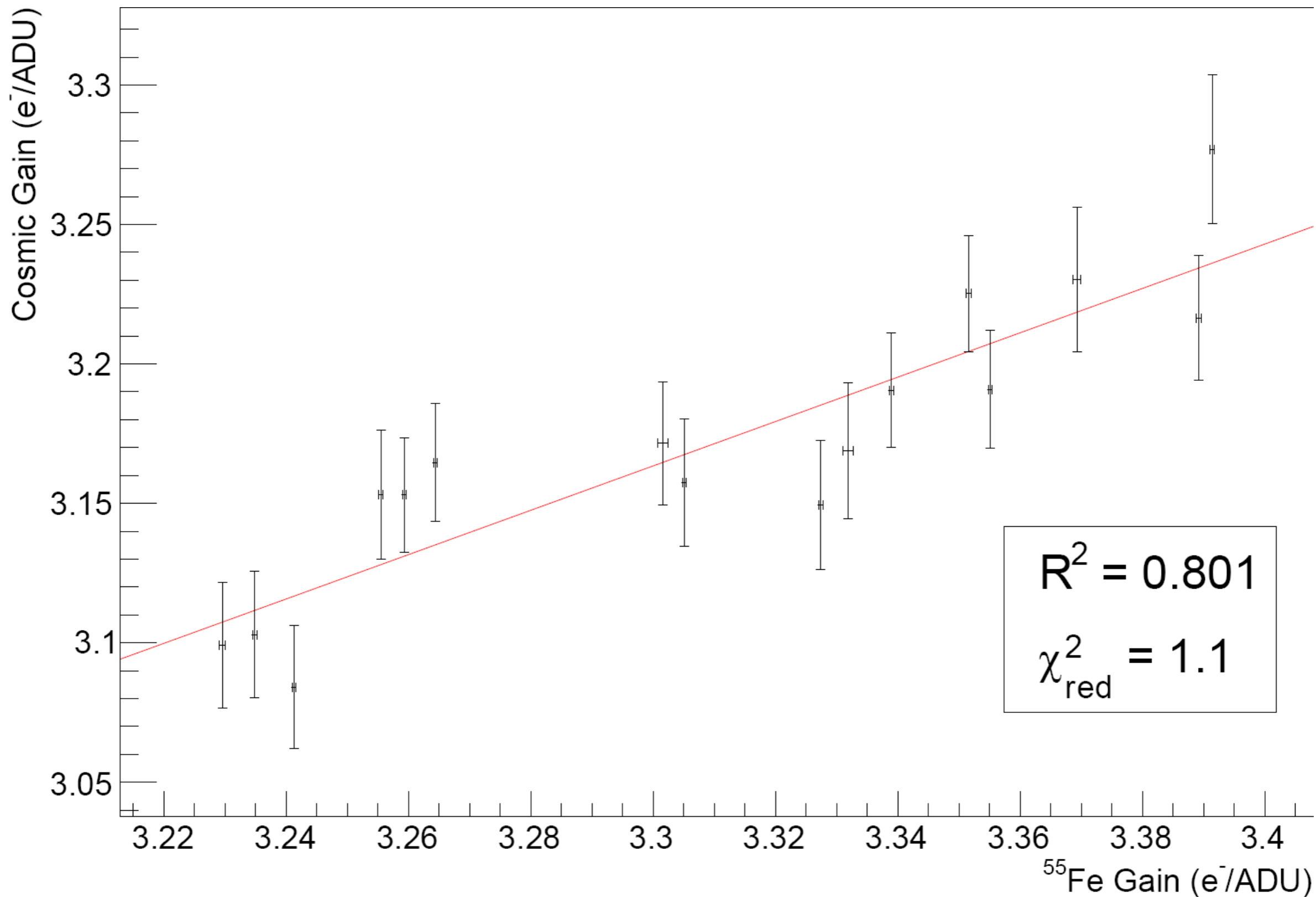
Entries = 38359

$\chi^2_{red} = 5.16$

Plot peaks against
each other for each
amplifier

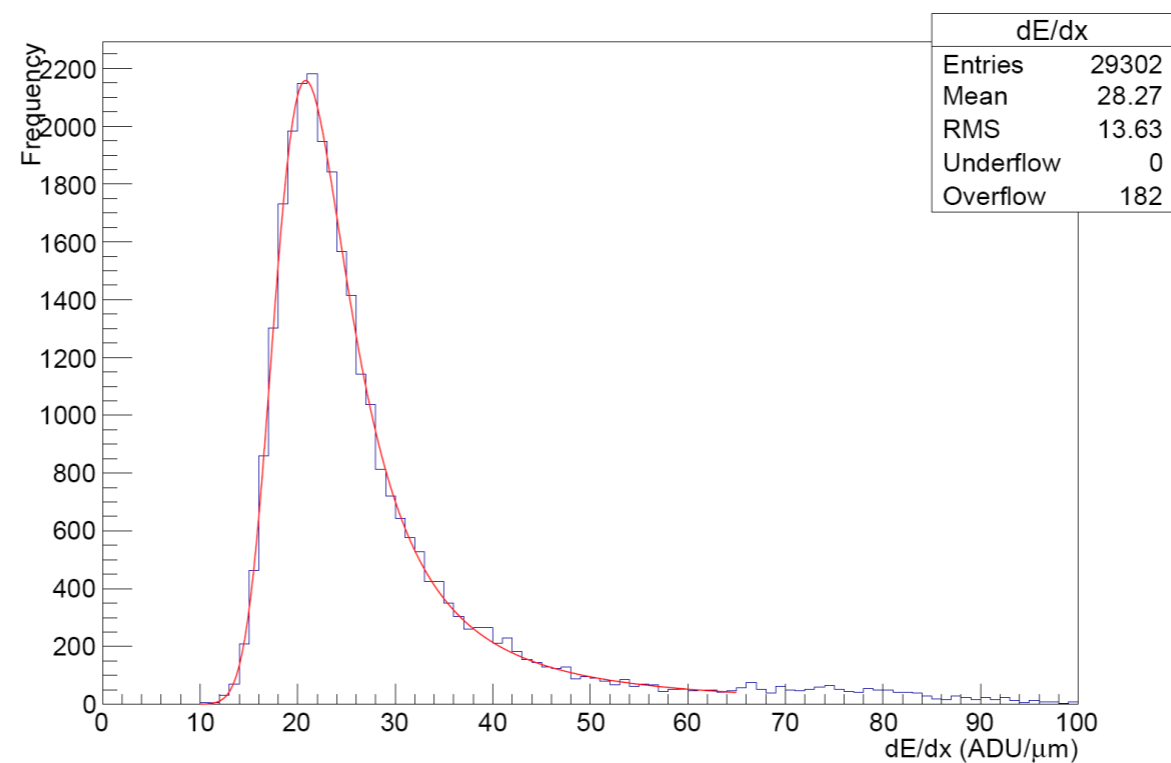


^{55}Fe vs cosmic gain measurement



▸ Conclusion:

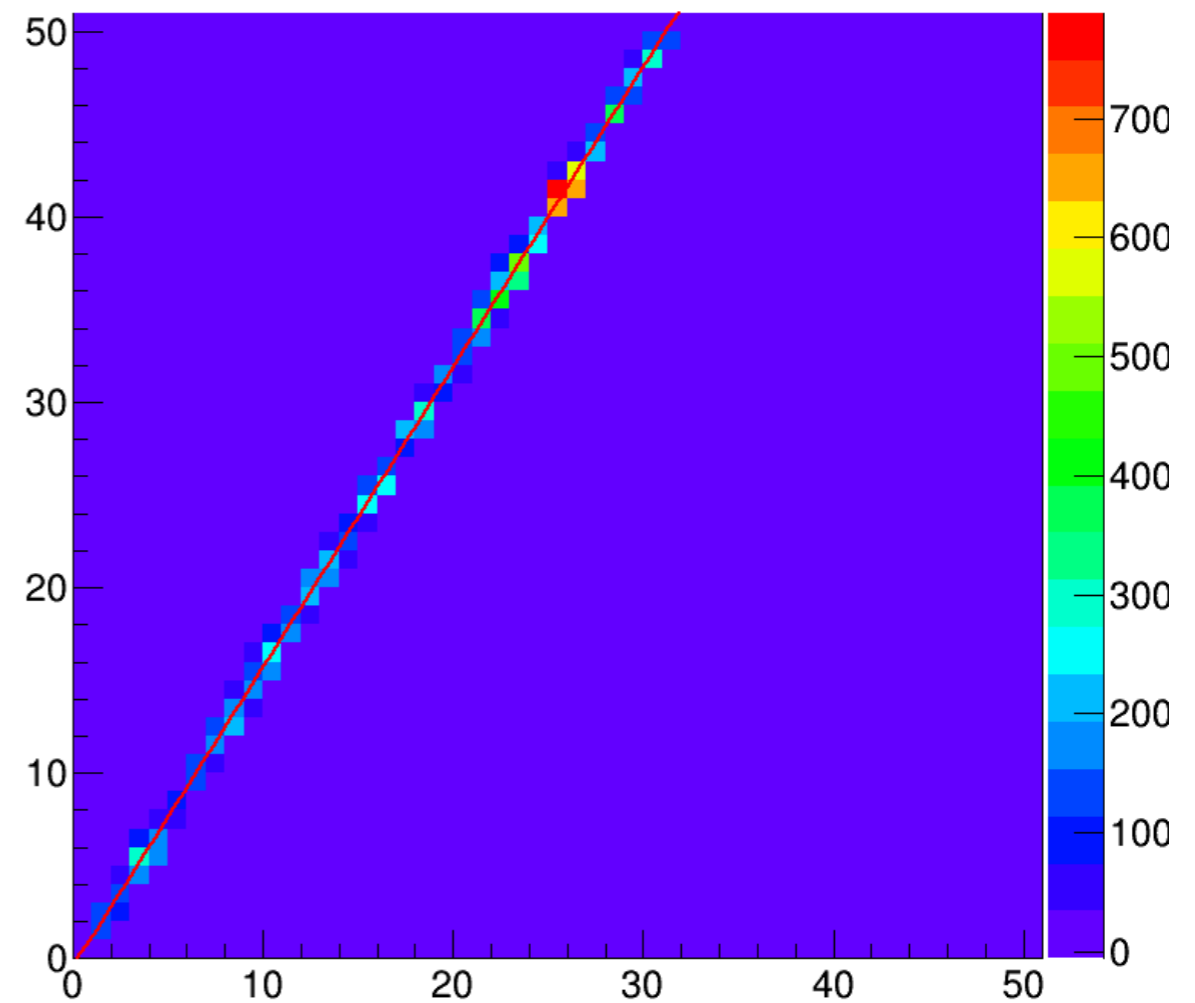
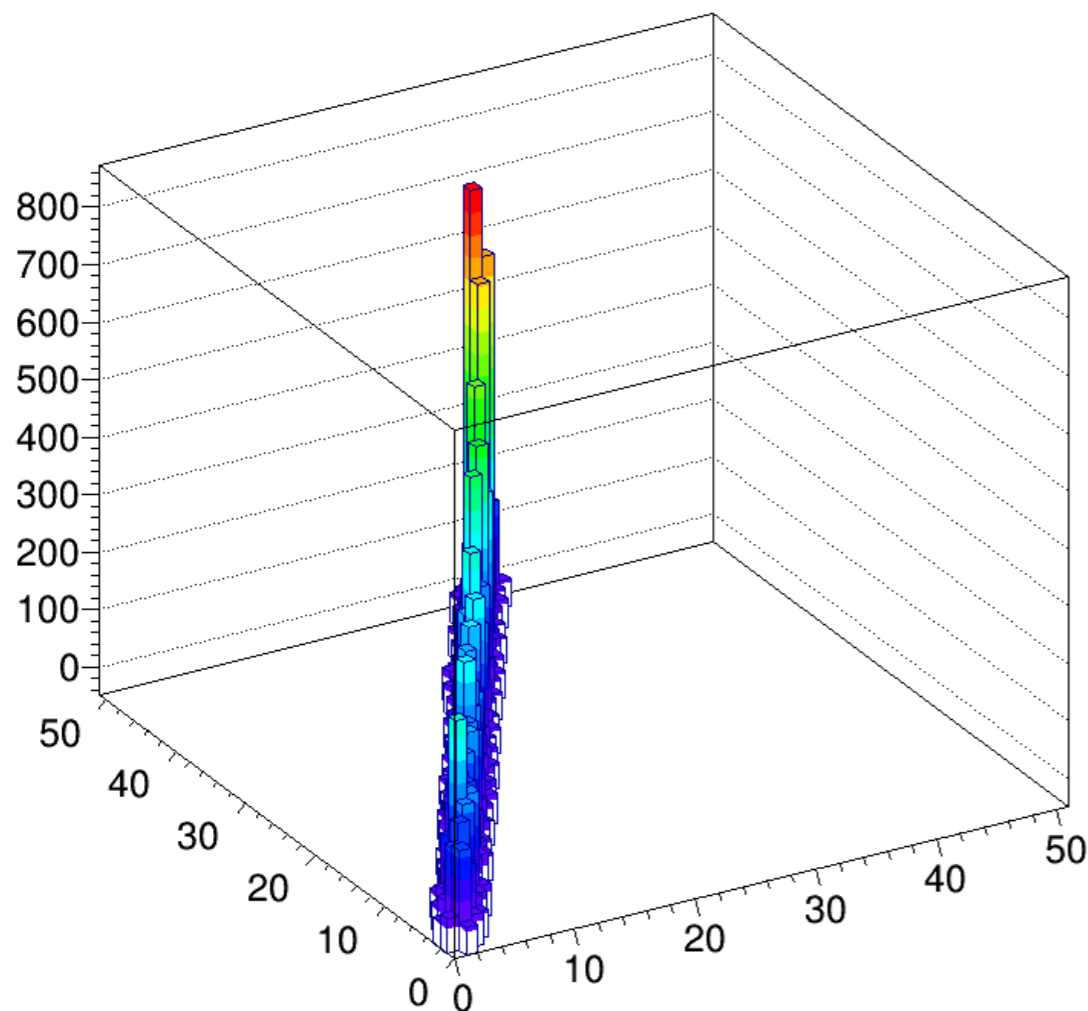
- You can measure gain without using ^{55}Fe or PTC
- You will get this information “for free” on the sky



Diffusion measurement

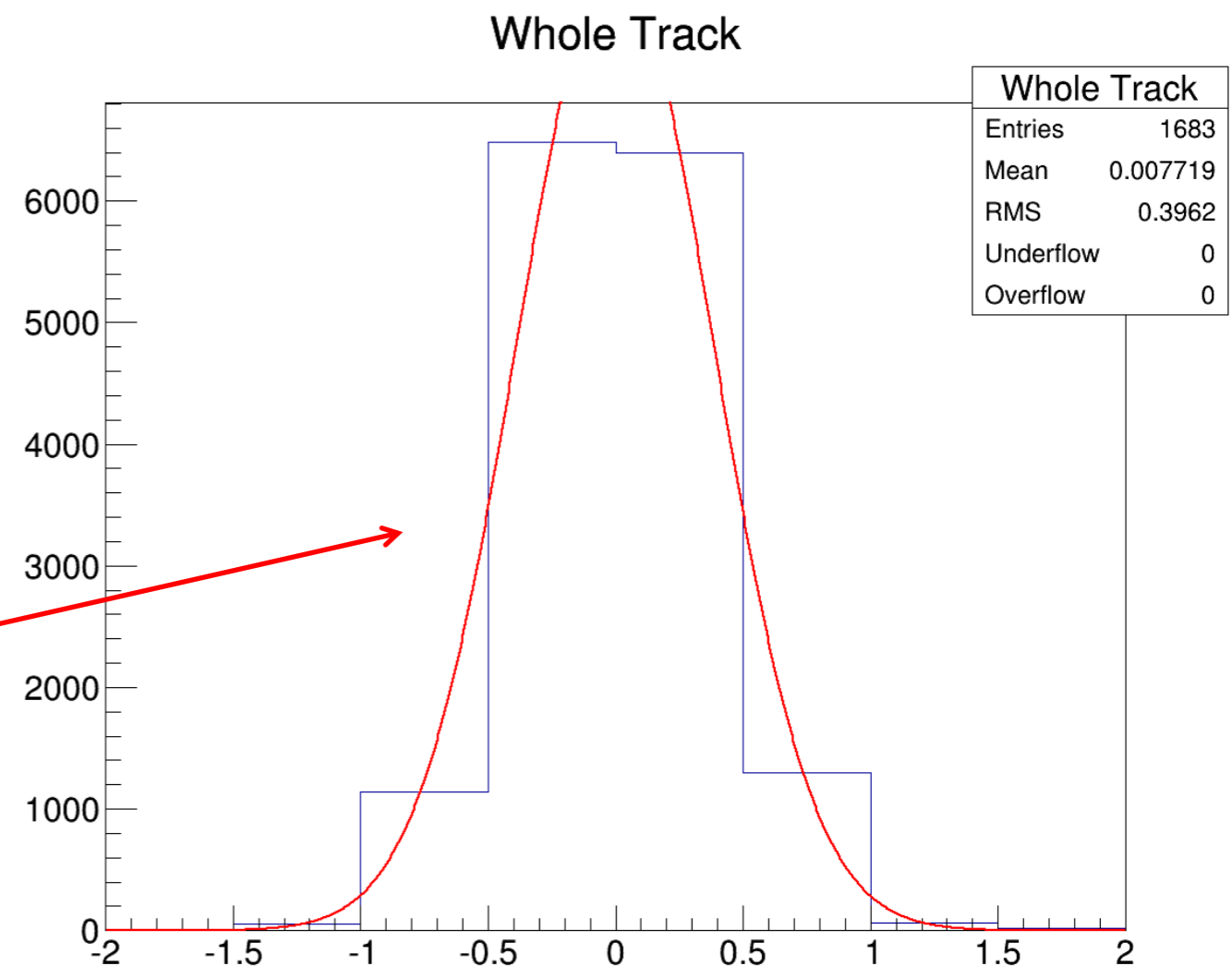
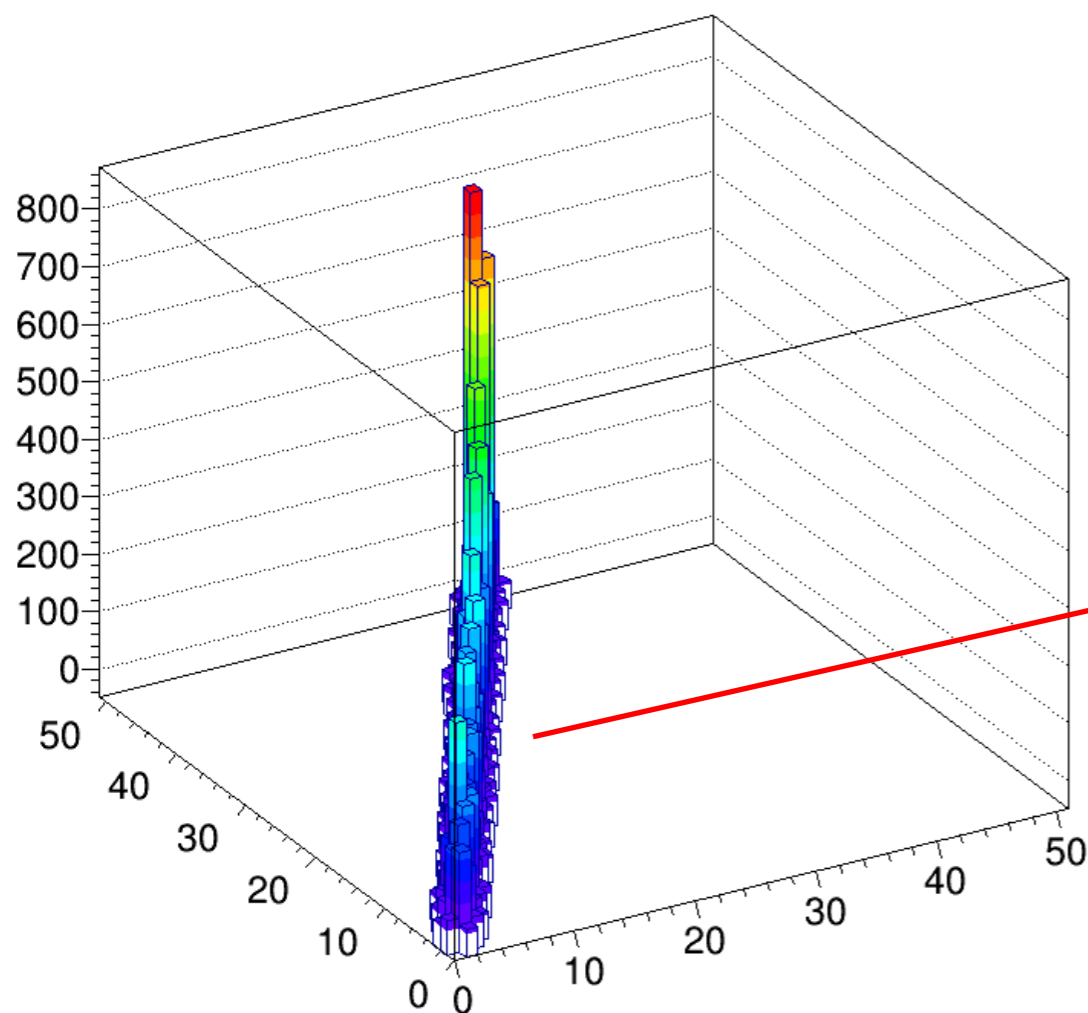
Diffusion Measurement

- Fit track with straight line
 - Plot intensity as a function of distance to the track
 - Fit a Gaussian to get the PSF



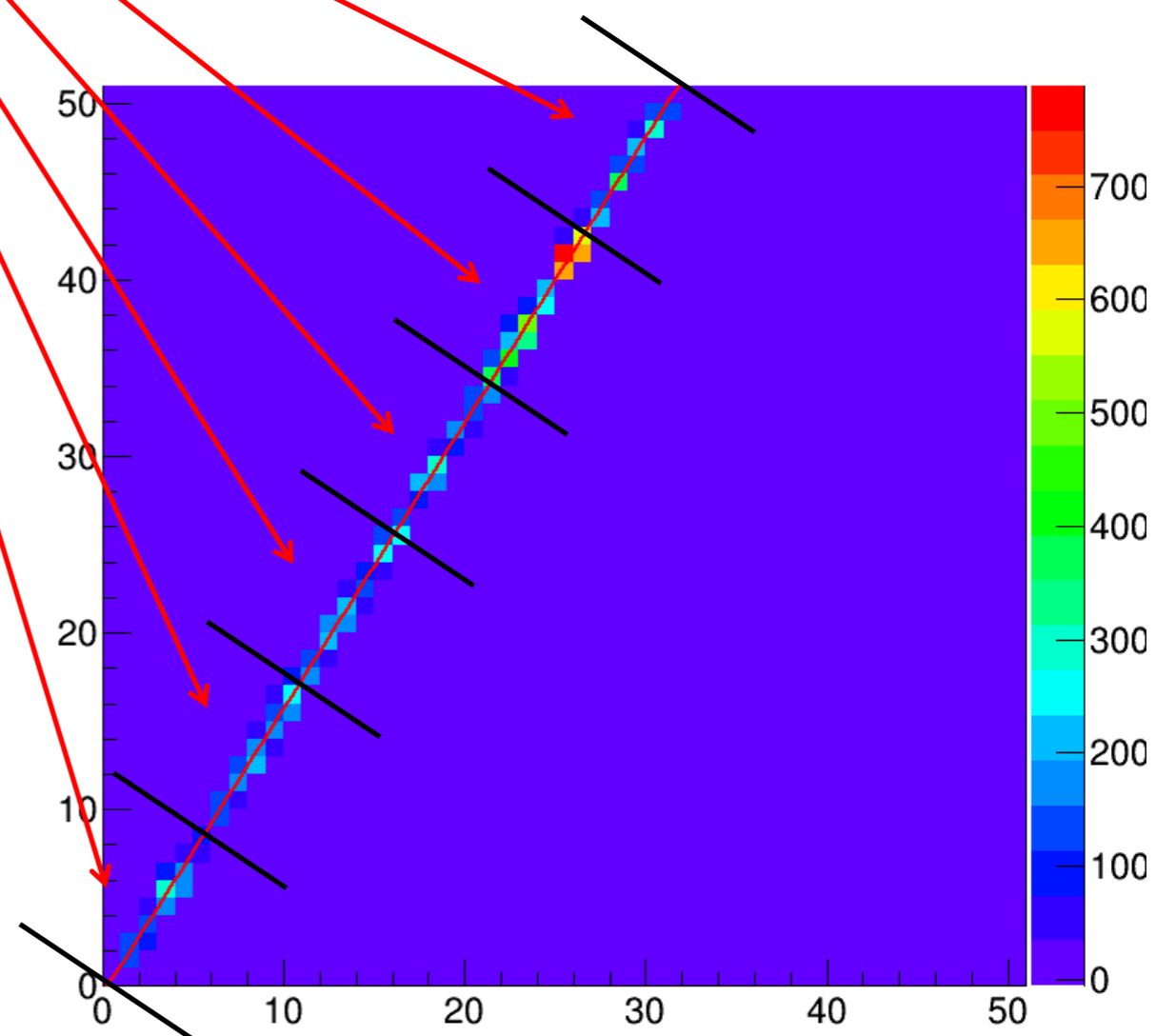
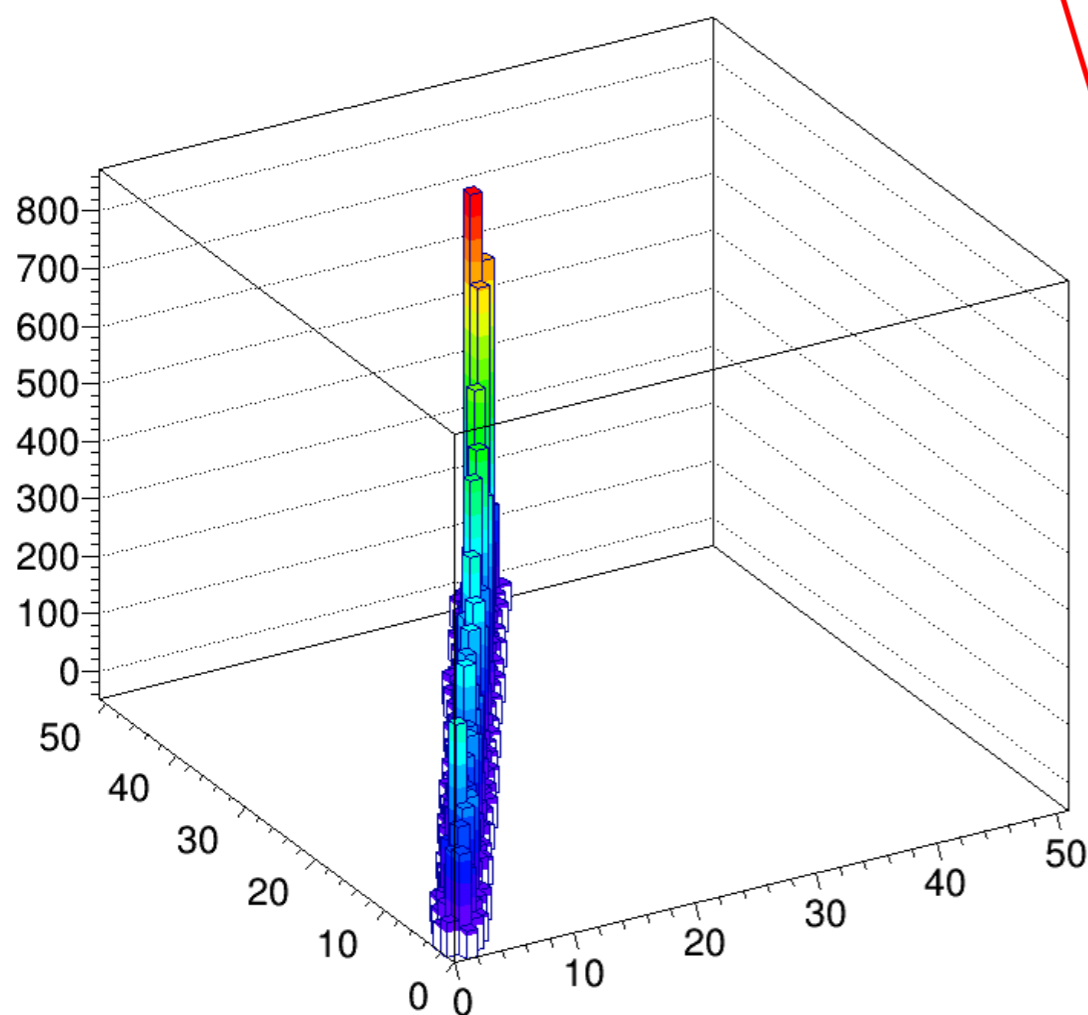
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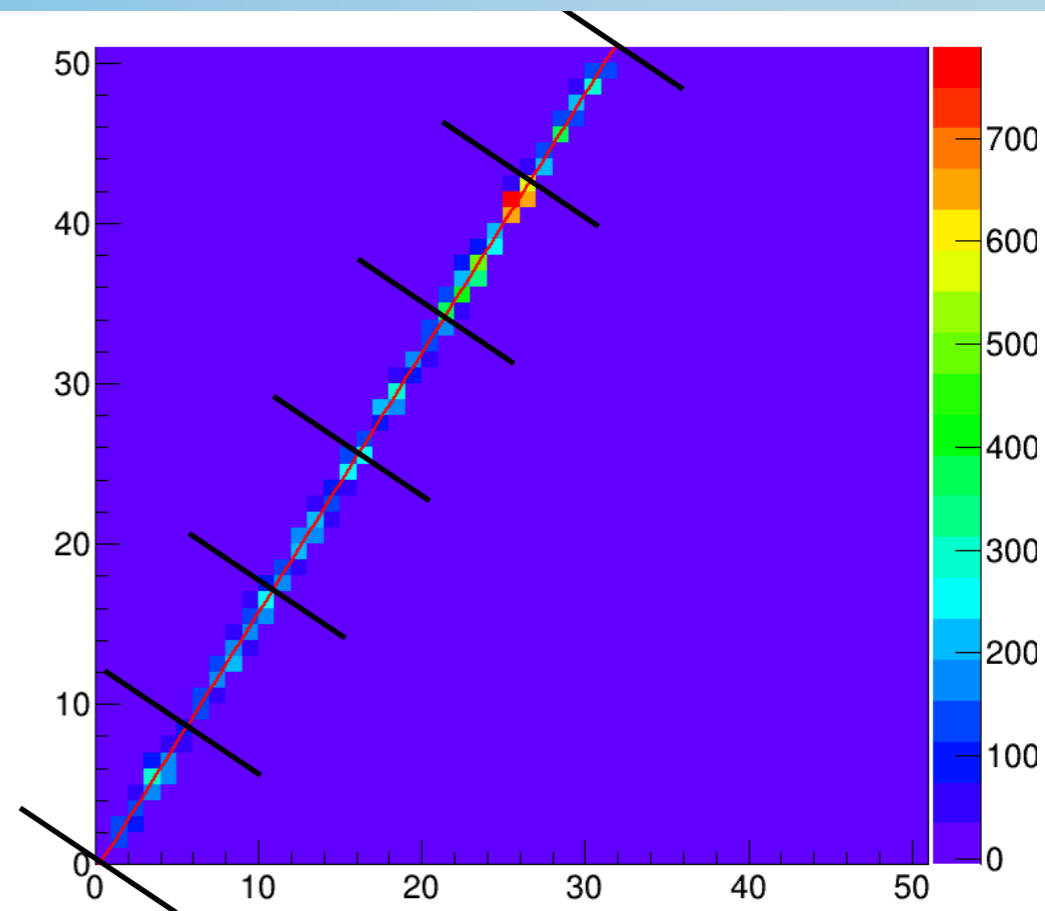
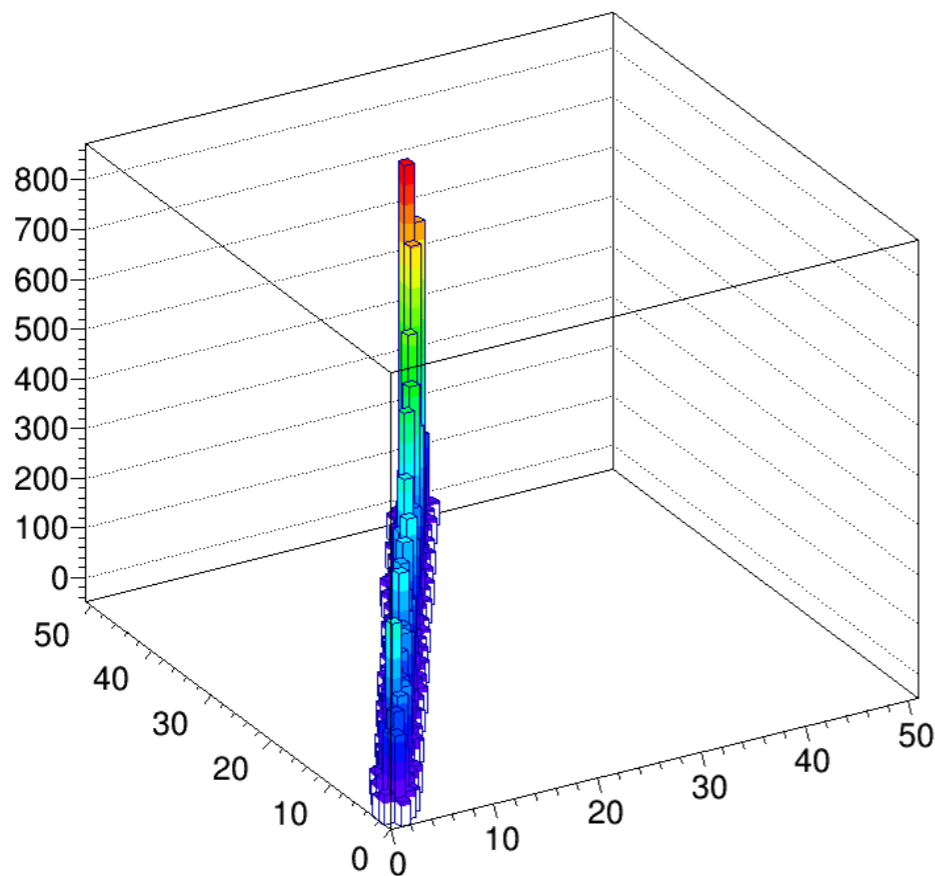


Track Segmentation

- To measure diffusion as a function of depth:
 - Divide track into n segments



Track Segmentation

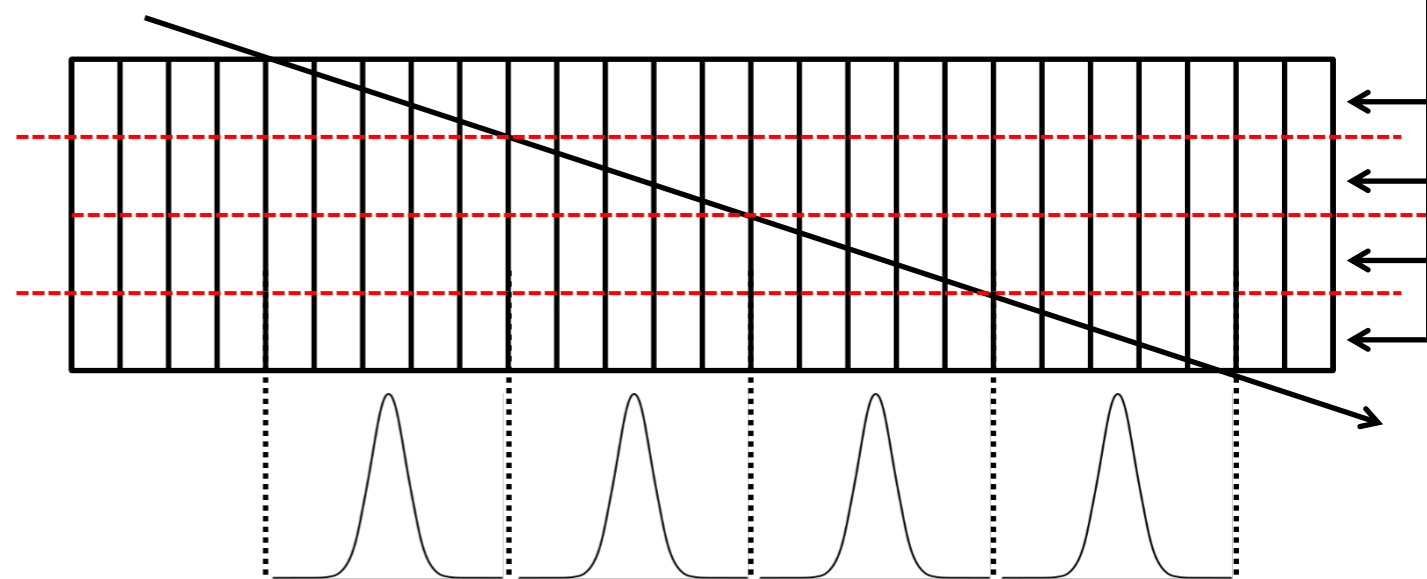


Take a long track

- ▶ Divide track into n segments
- ▶ For each segment:
 - ▶ Produce histogram of intensity as a function of distance to the track
 - ▶ Fit a Gaussian to get the PSF
- ▶ Plot the widths of these Gaussians w.r.t. track segment number (i.e. depth)

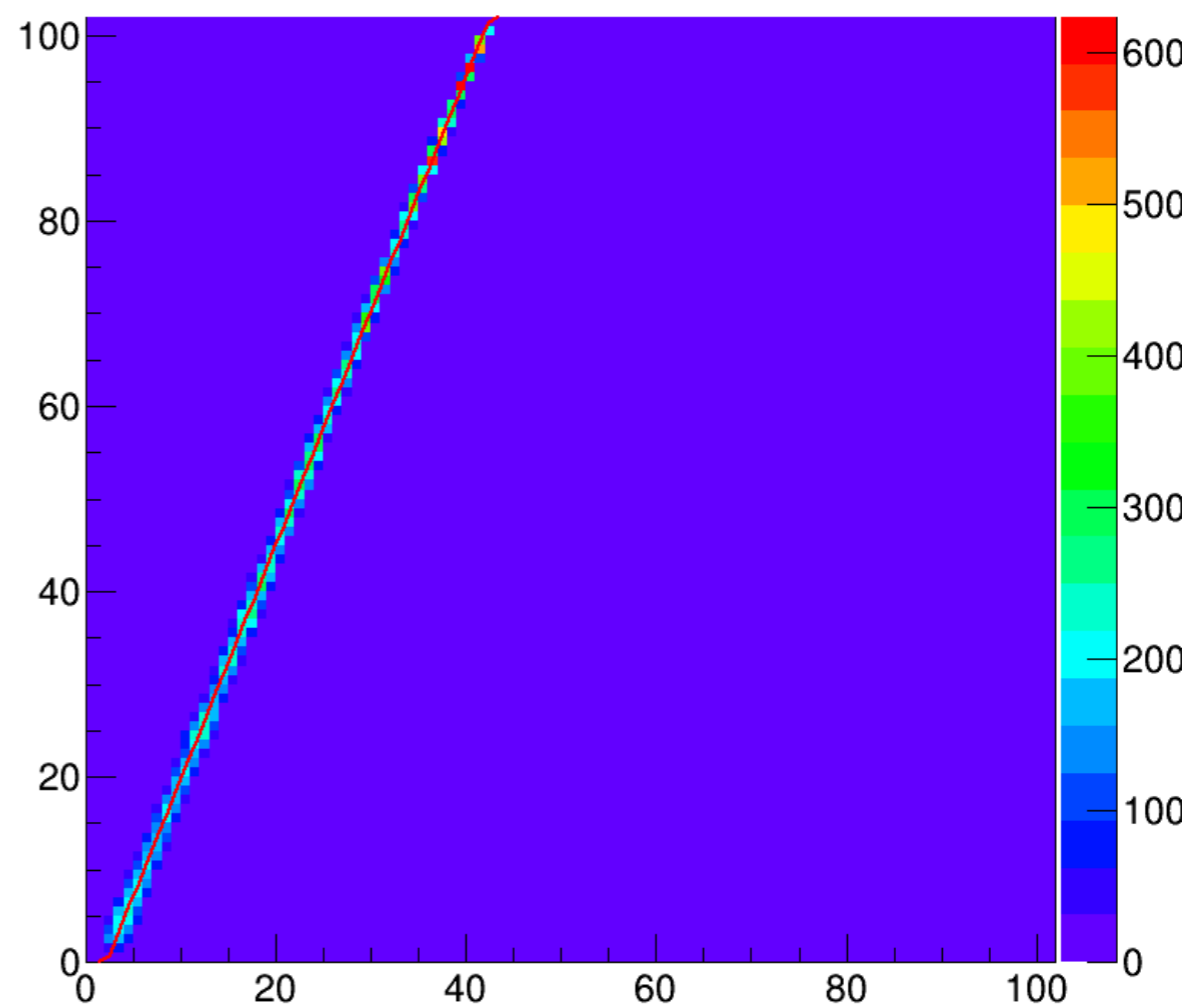
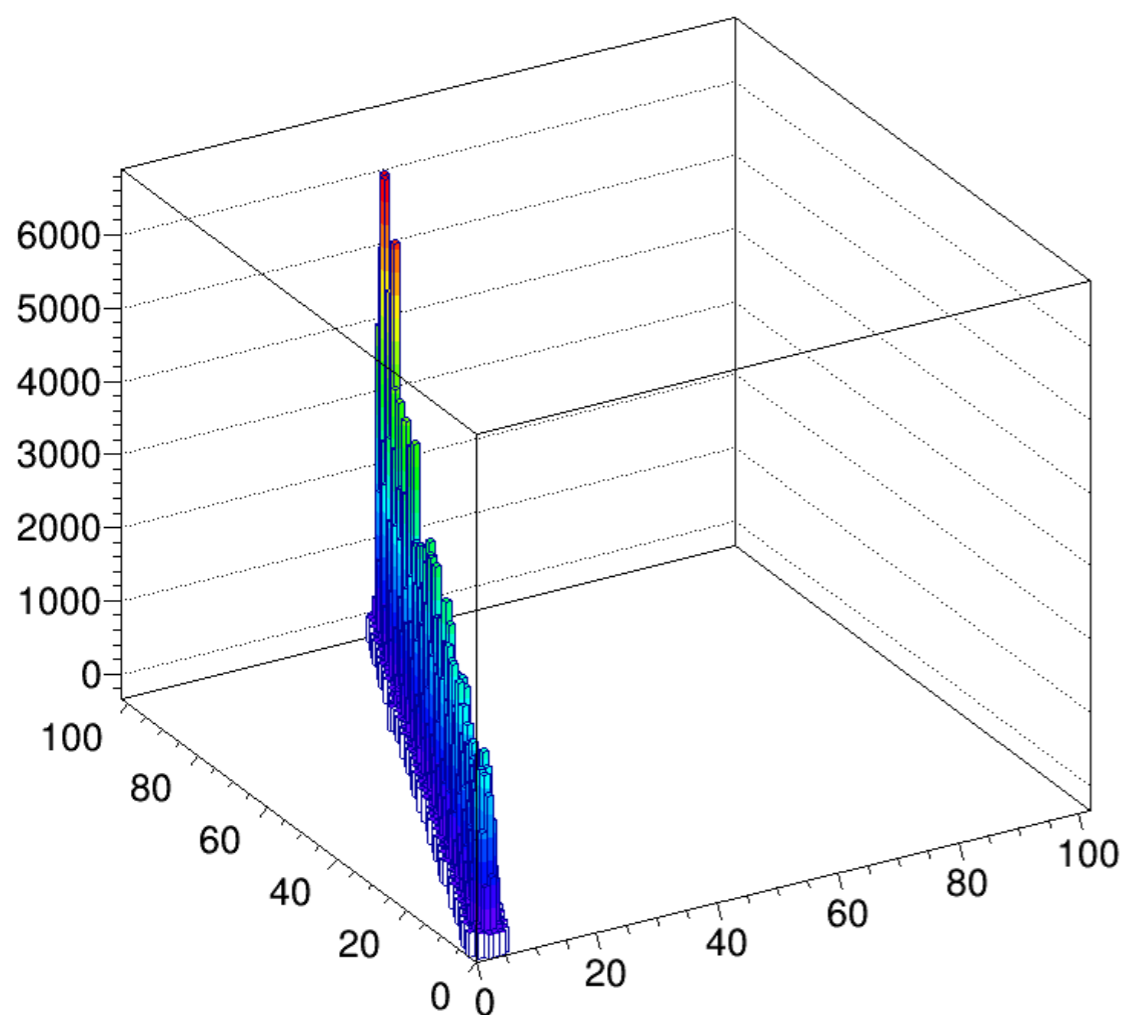
Long muon track

average collection depths



Track Segmentation

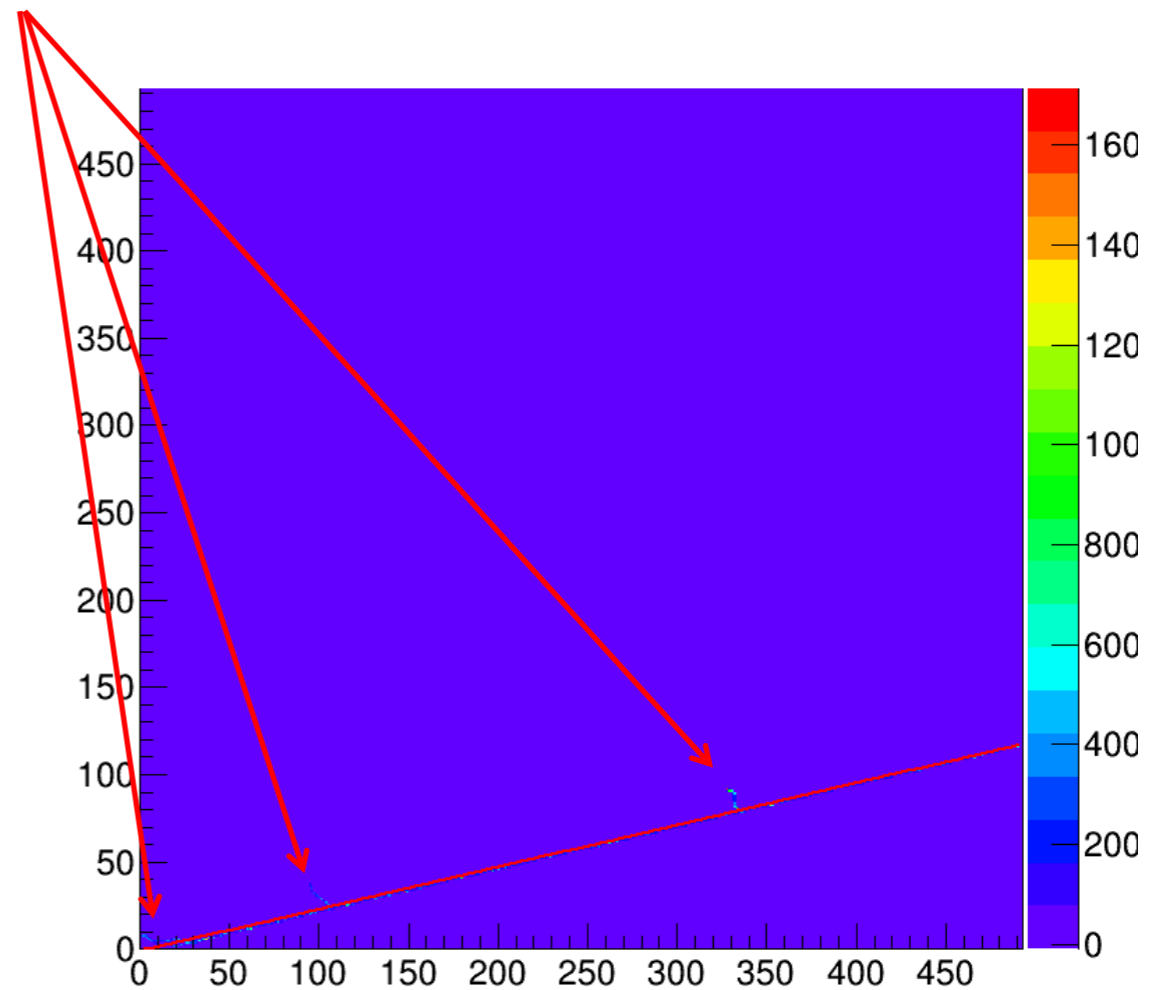
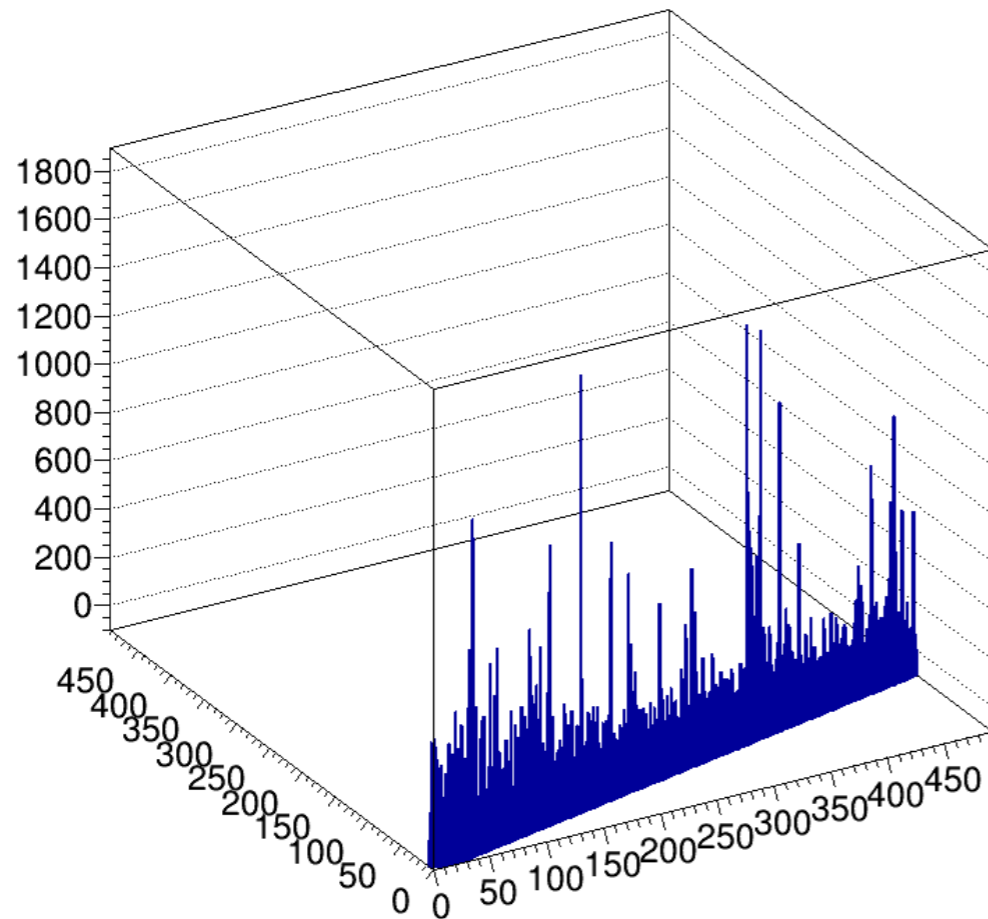
- Nice example of a track changing in PSF with depth in sensor



Track cleaning

- ▶ However, unlike for gain measurement, when measuring diffusion δ -rays must be *excluded*
 - ▶ They will ruin the PSF measurement for the depth at which they occur

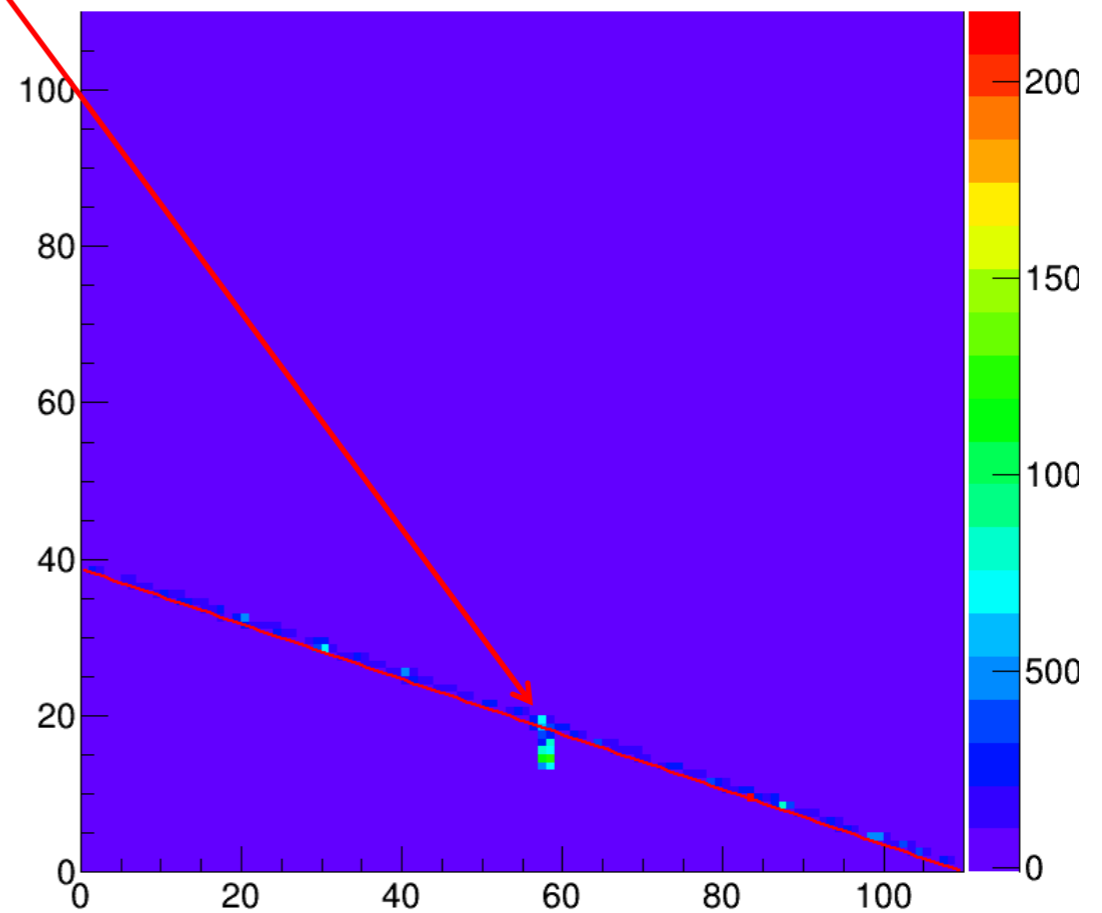
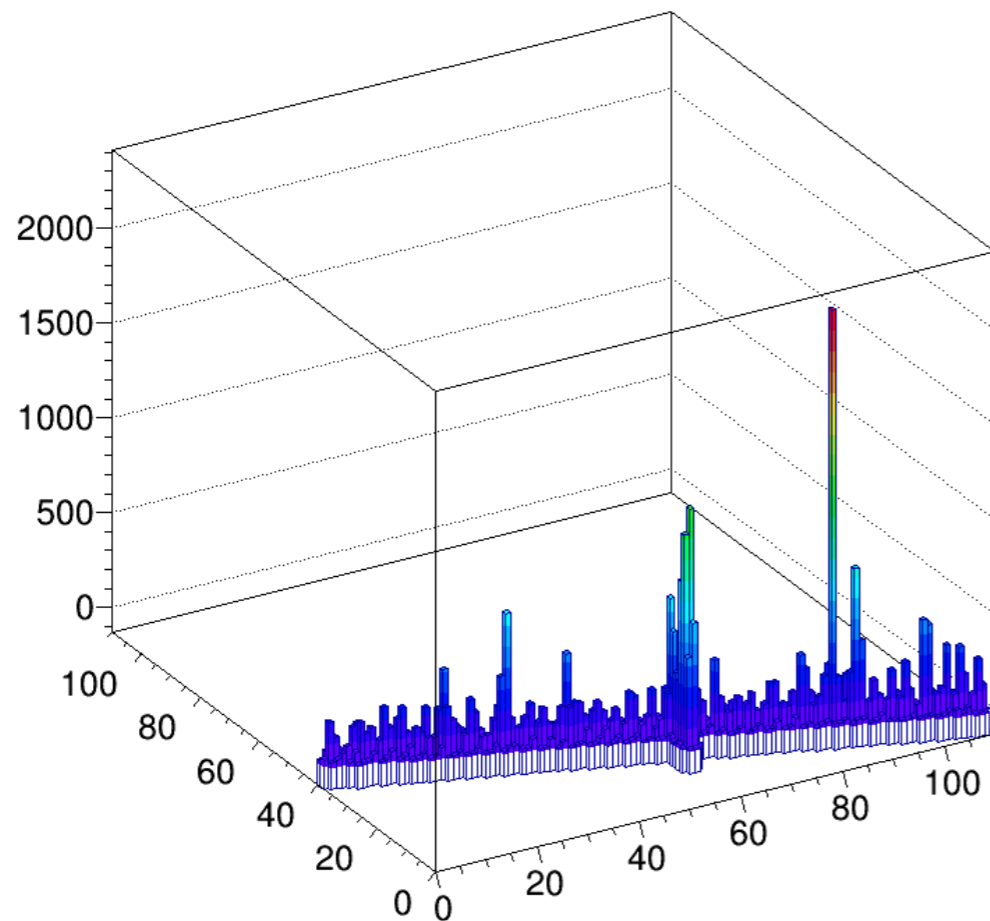
Remove tracks with deltas



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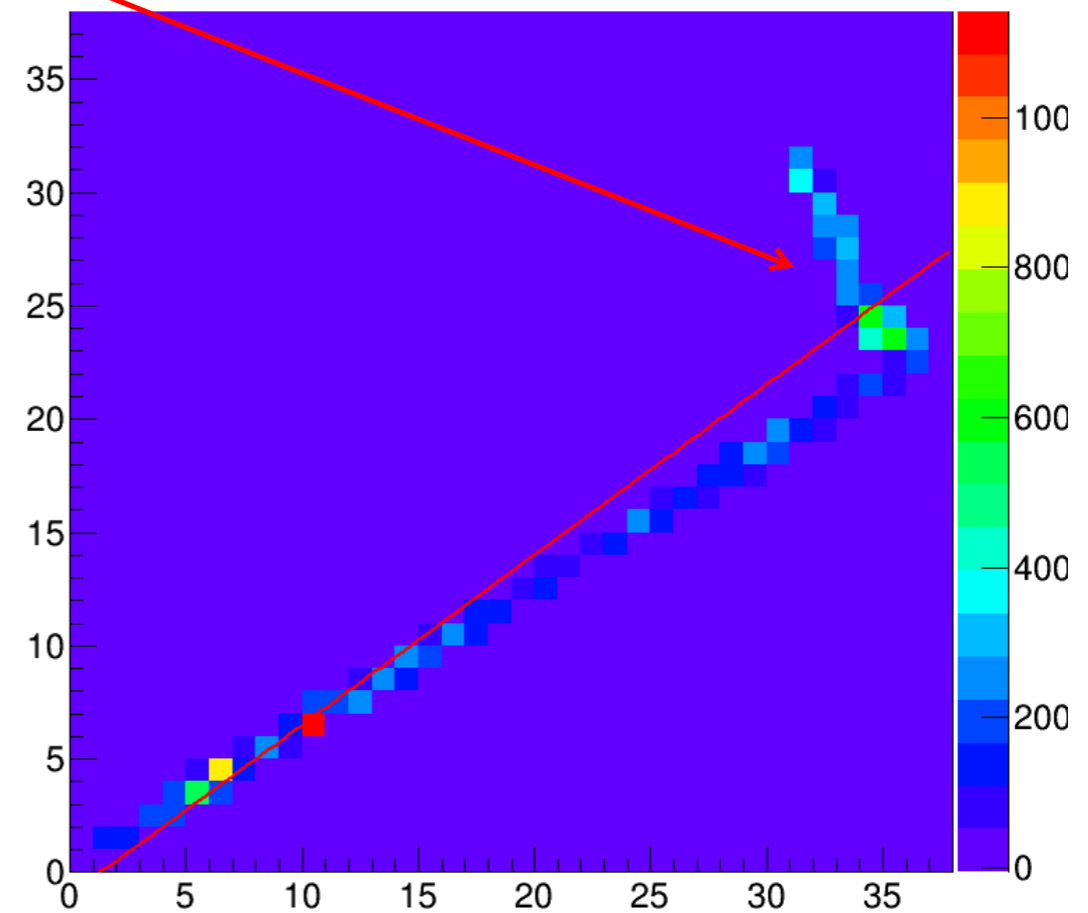
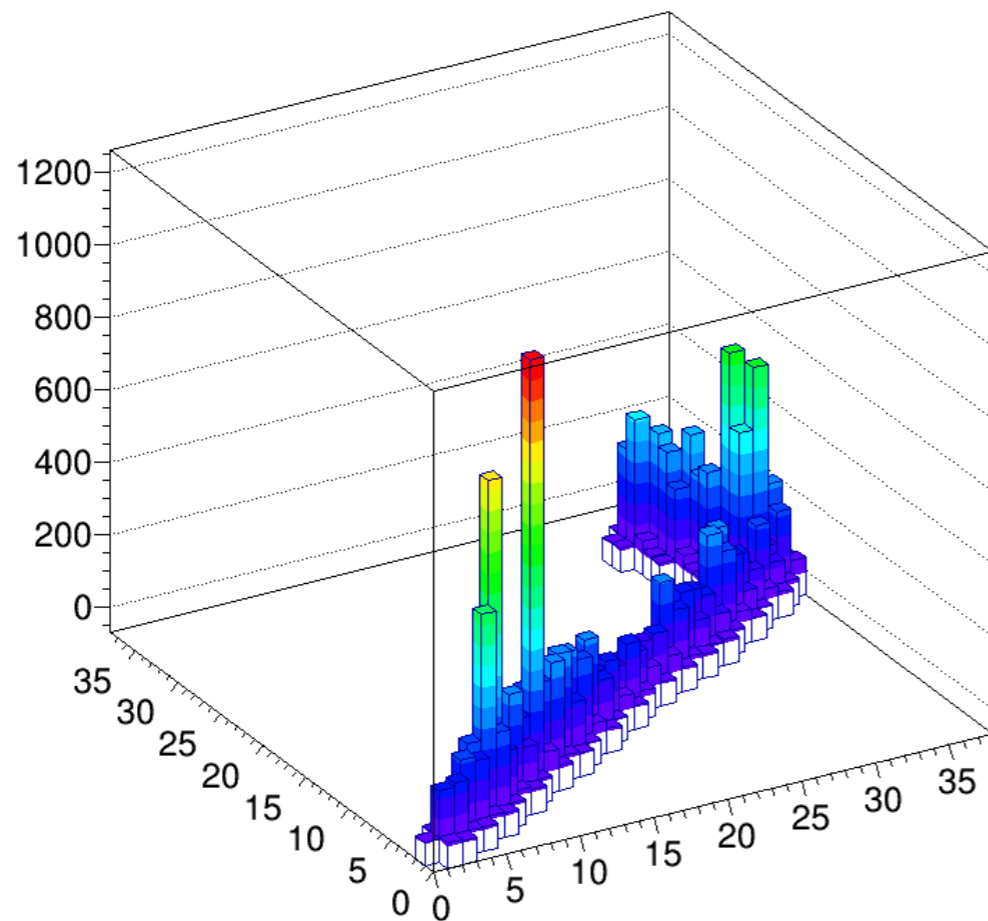
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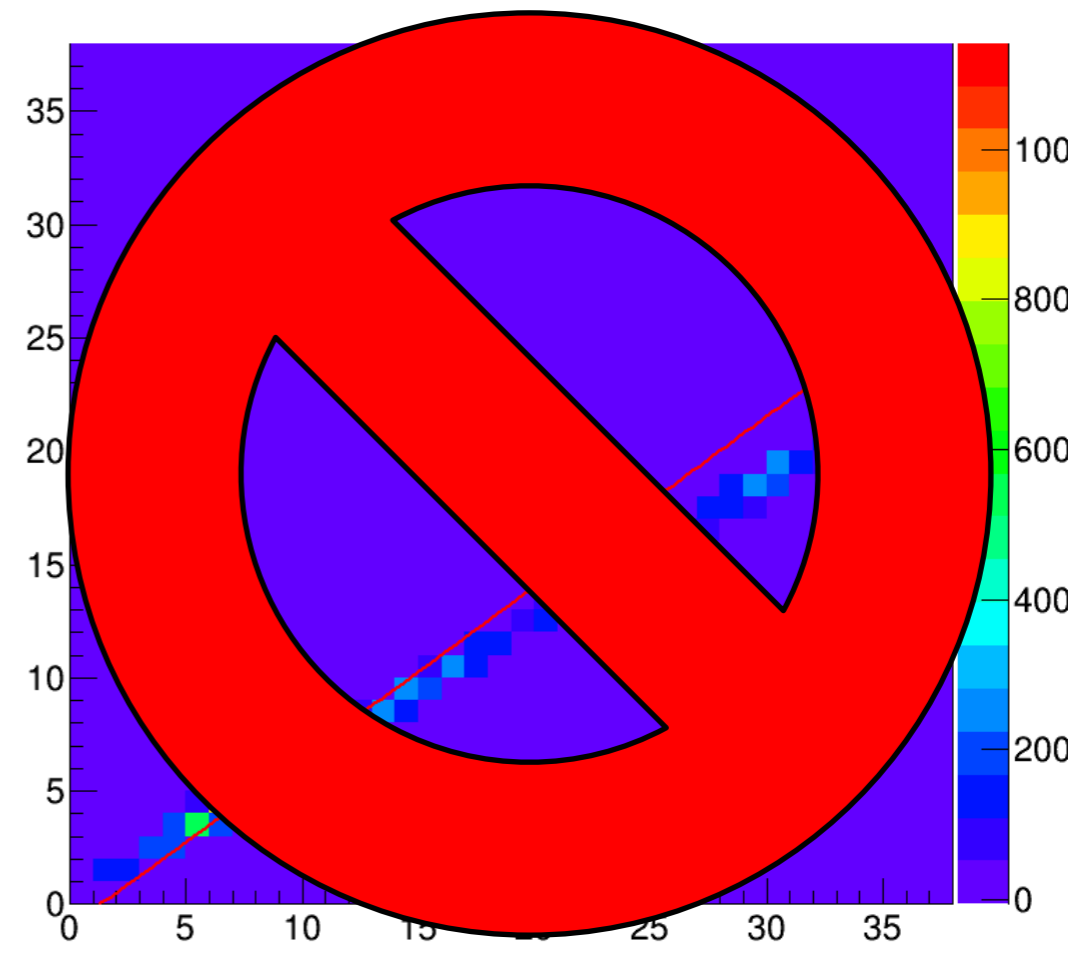
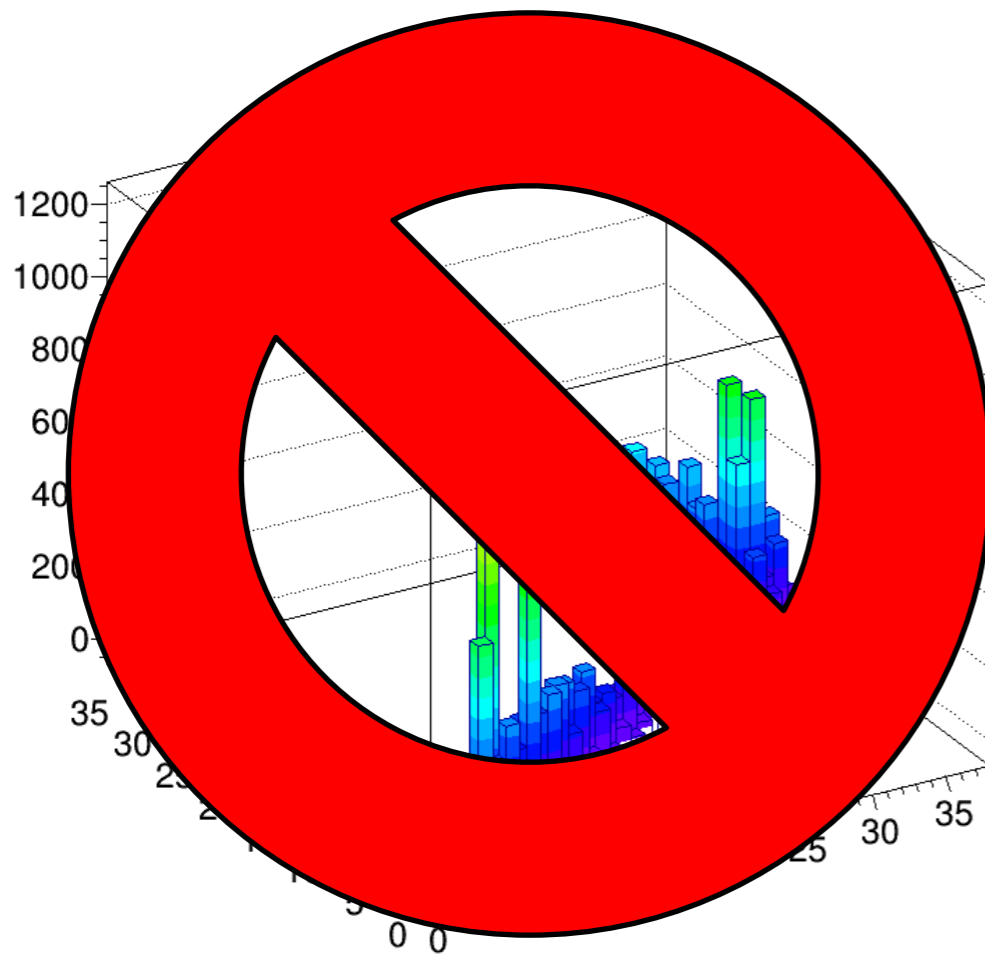
Remove tracks with deltas



Track cleaning

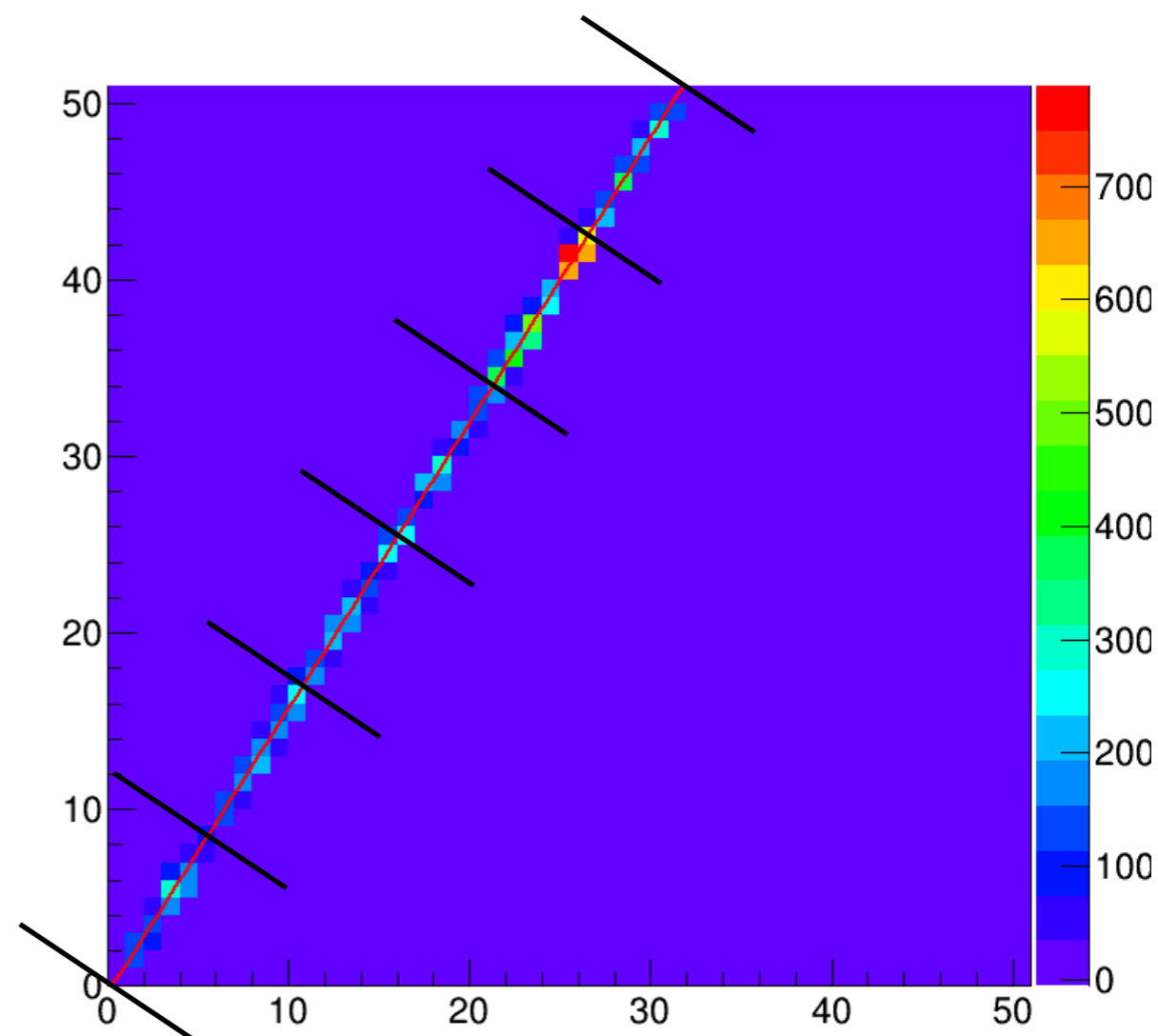
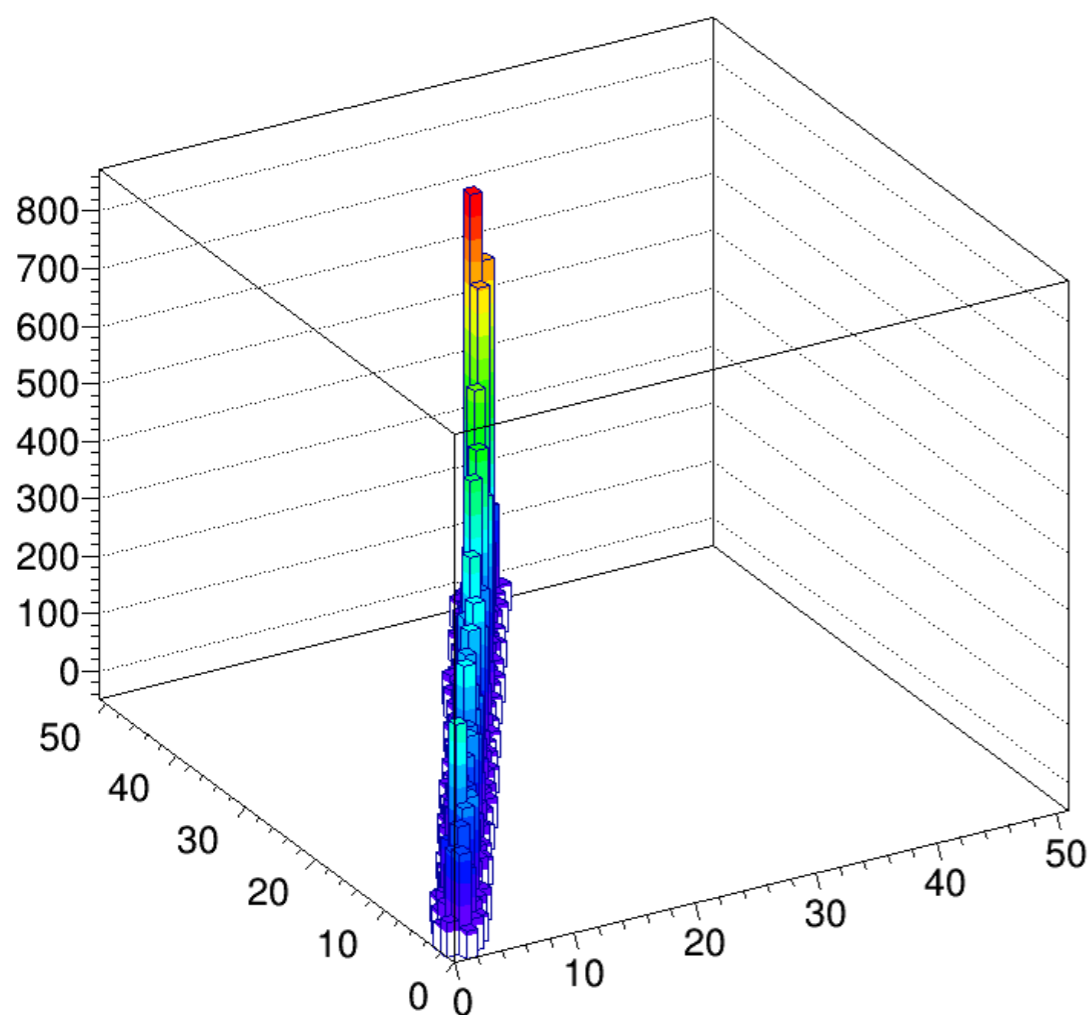
- Find all tracks in dataset
 - Cut dataset to only include long & straight tracks
 - Remove tracks with large delta-rays as these ruin PSF for a certain depth

Remove tracks with deltas



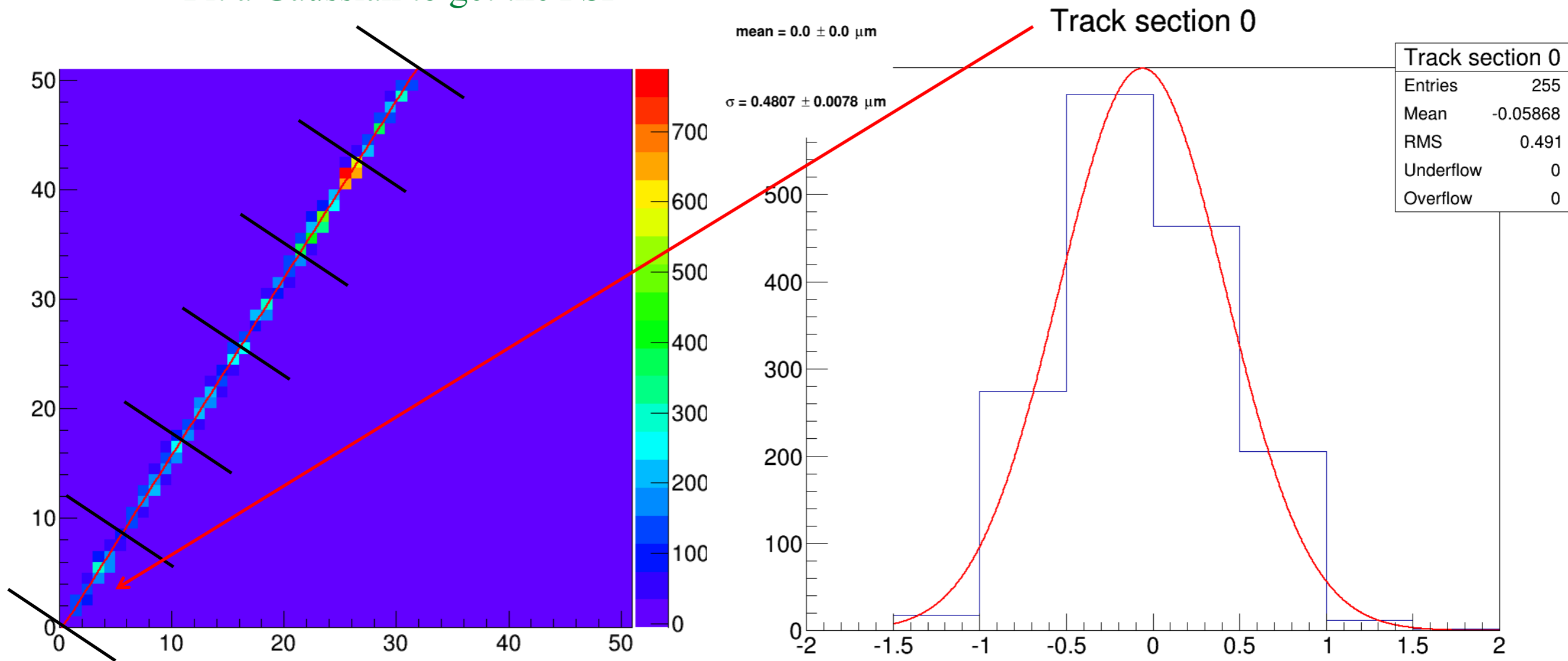
Track Segmentation

- Fit each track with straight line
 - Divide track into n segments
 - For each segment:
 - Produce histogram of intensity as a function of distance to the track
 - Fit a Gaussian to get the PSF



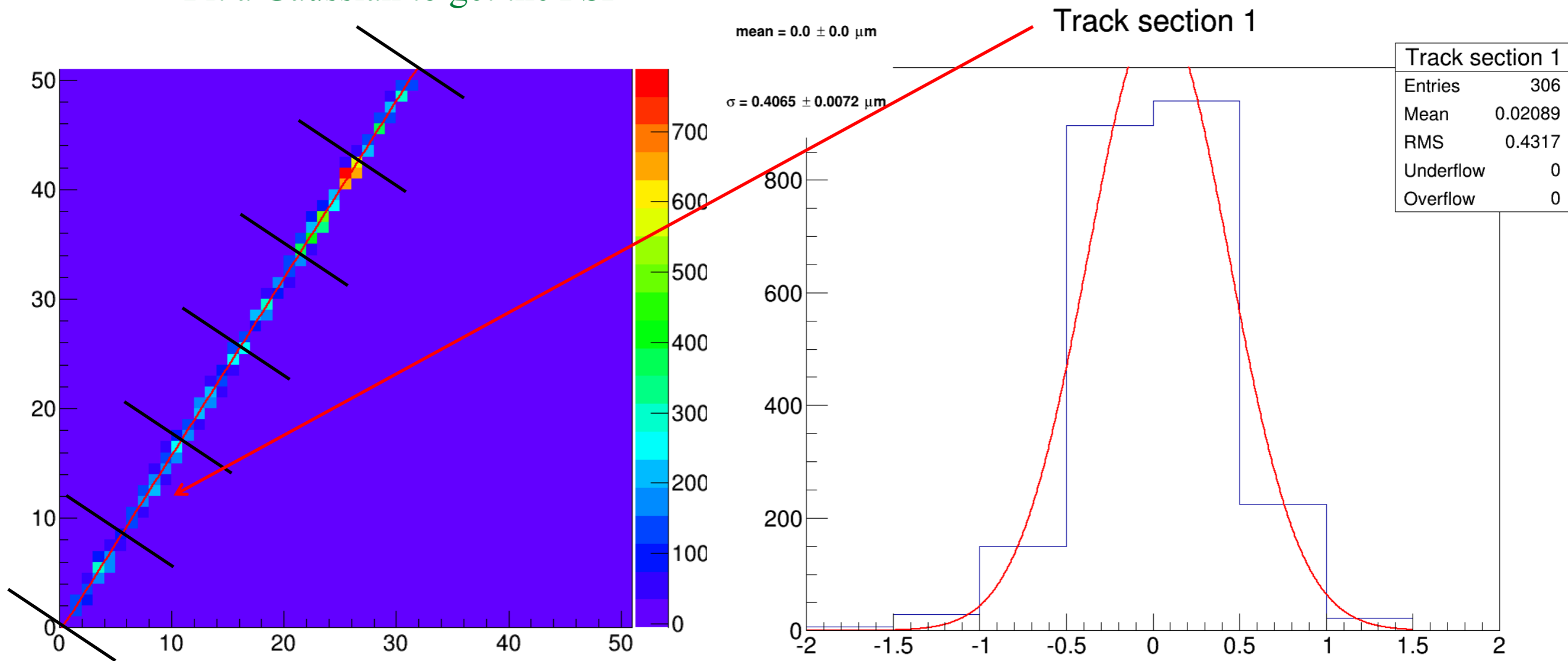
Segment Measurement

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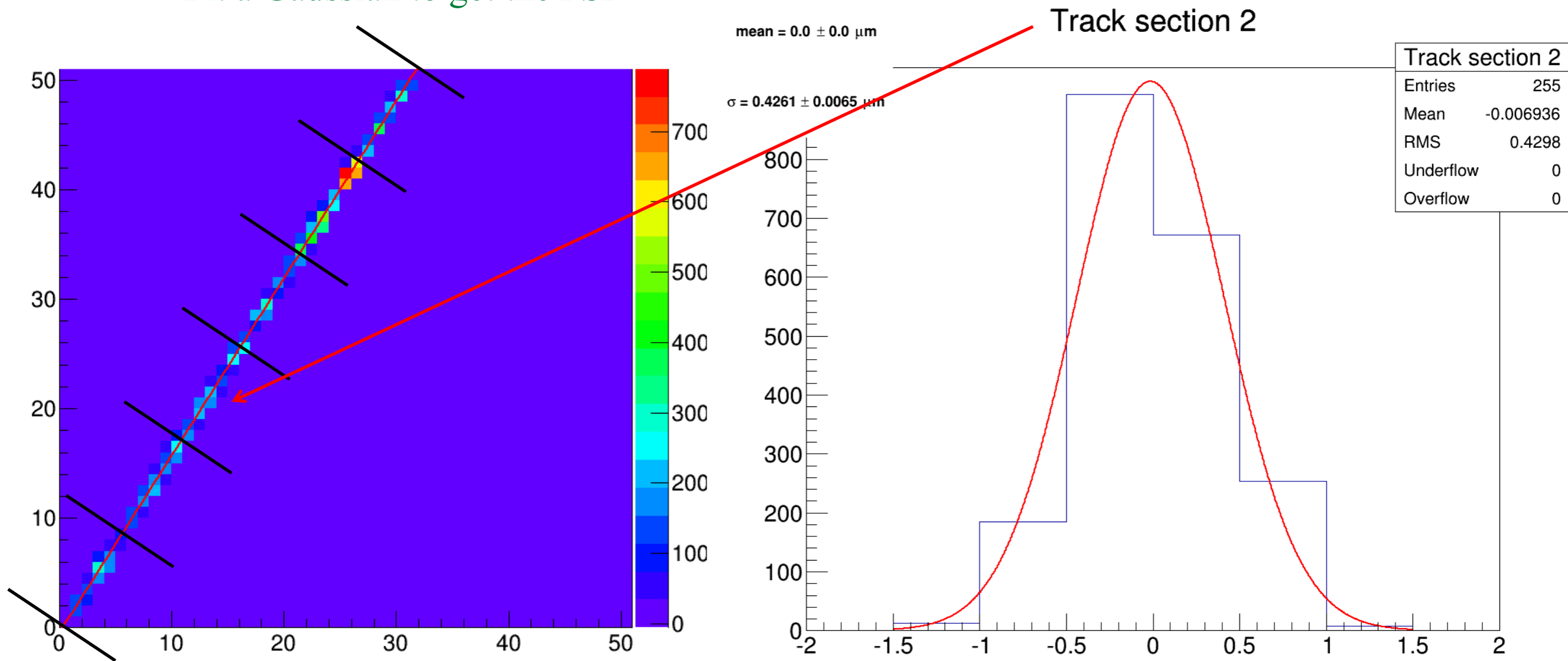
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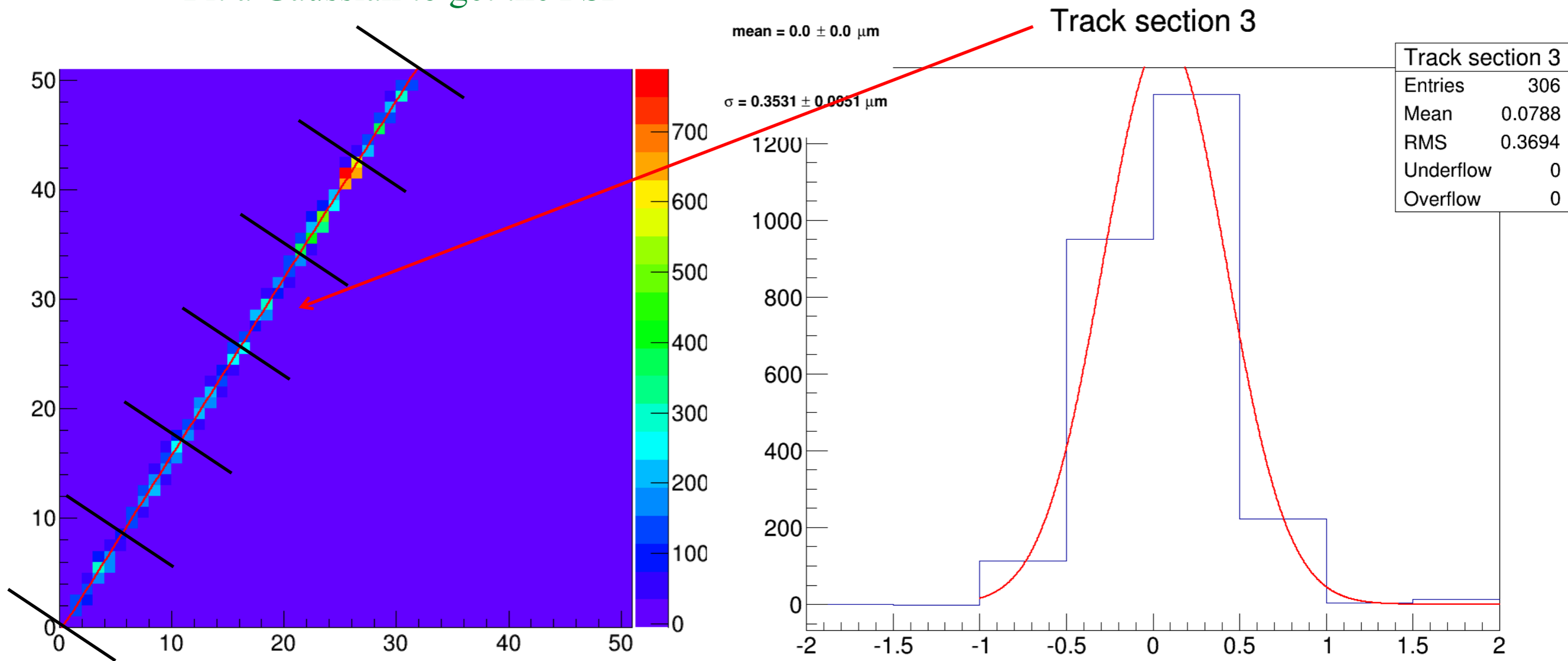
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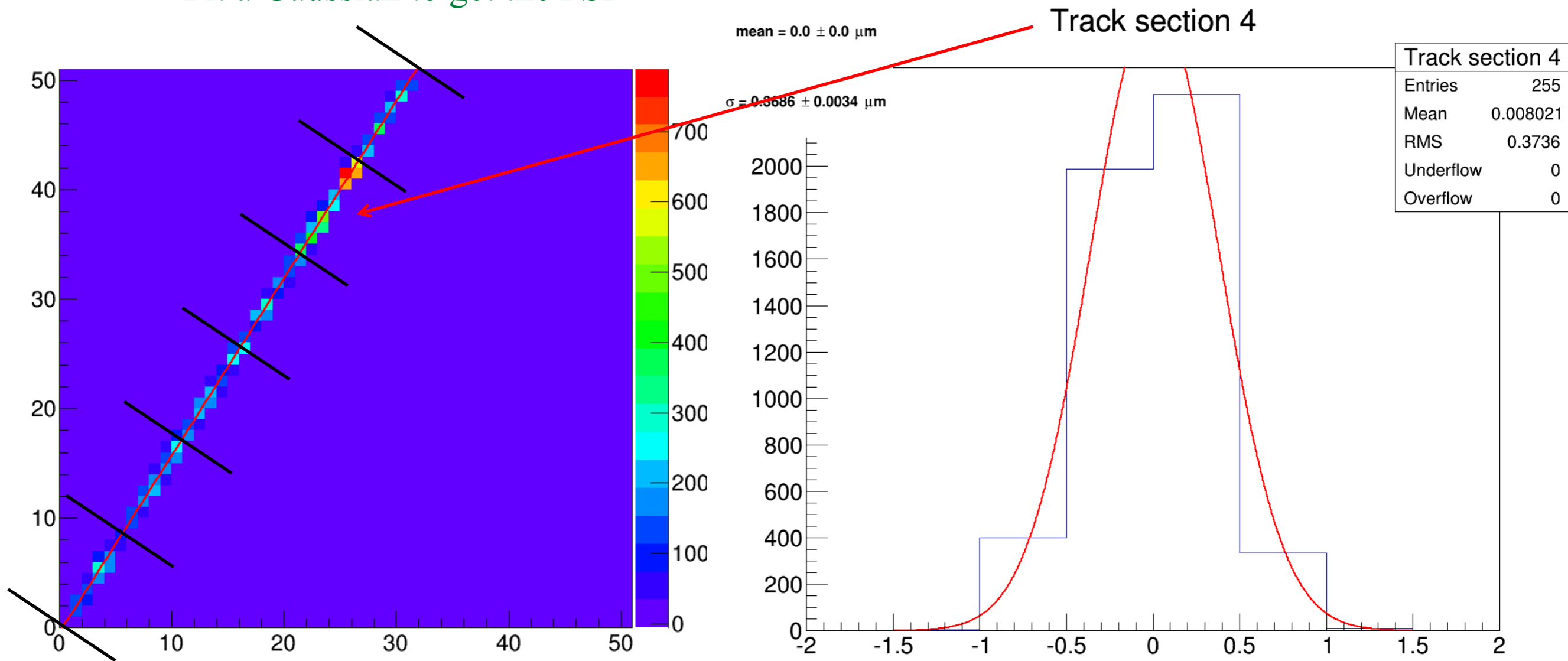
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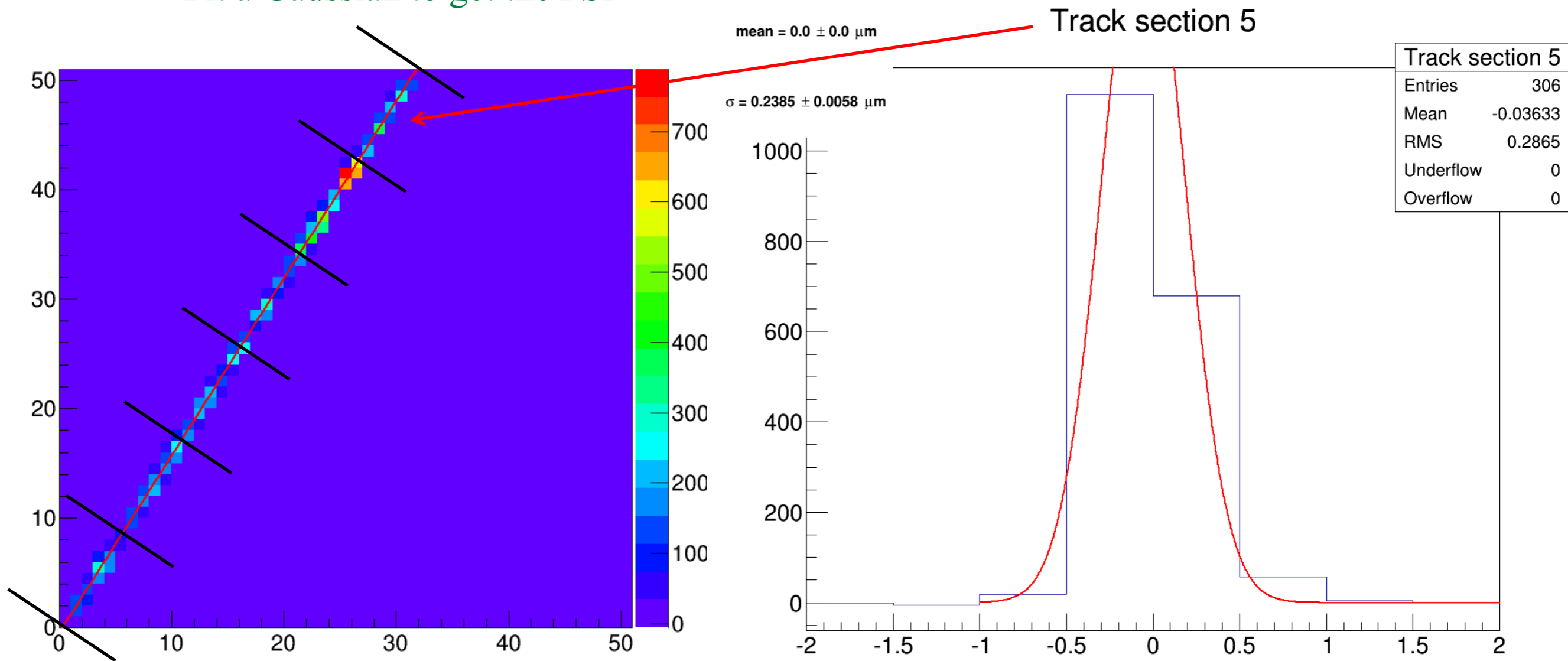
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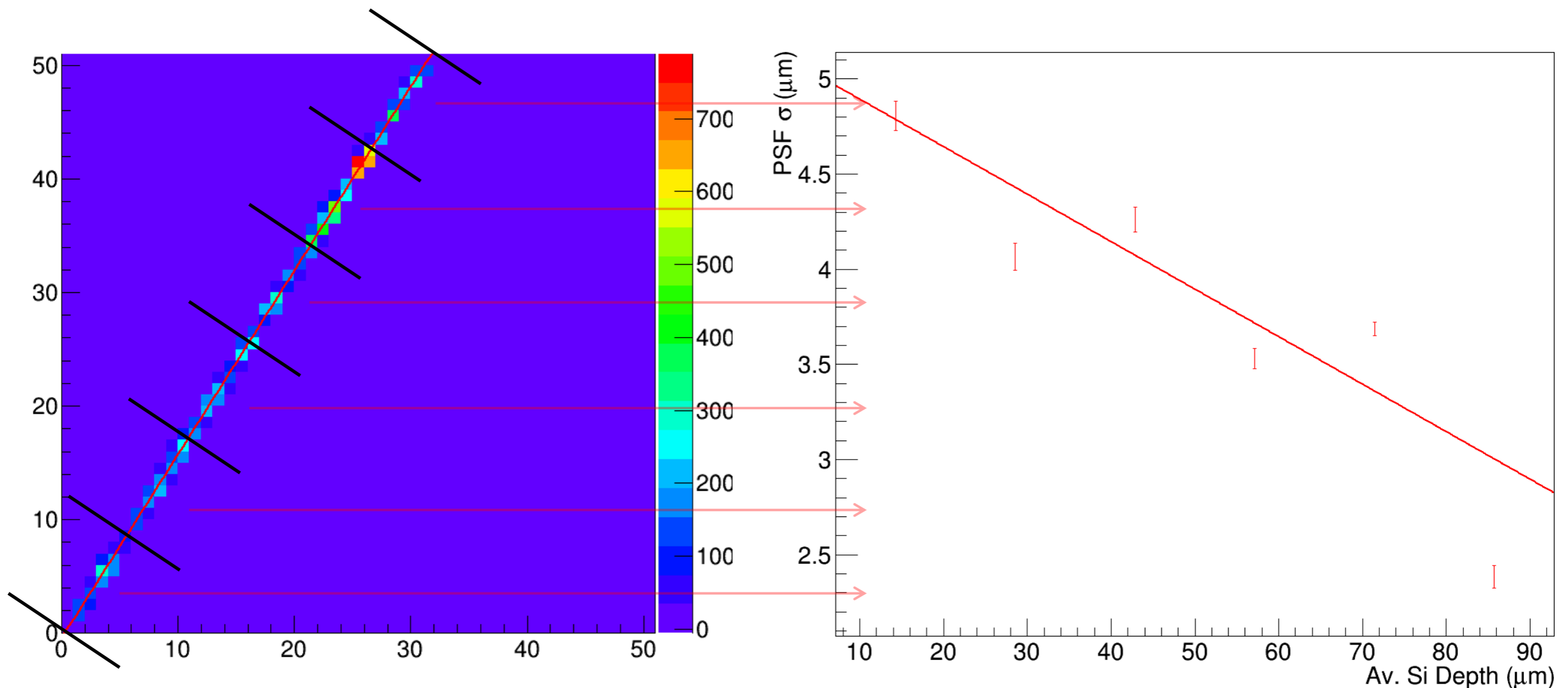
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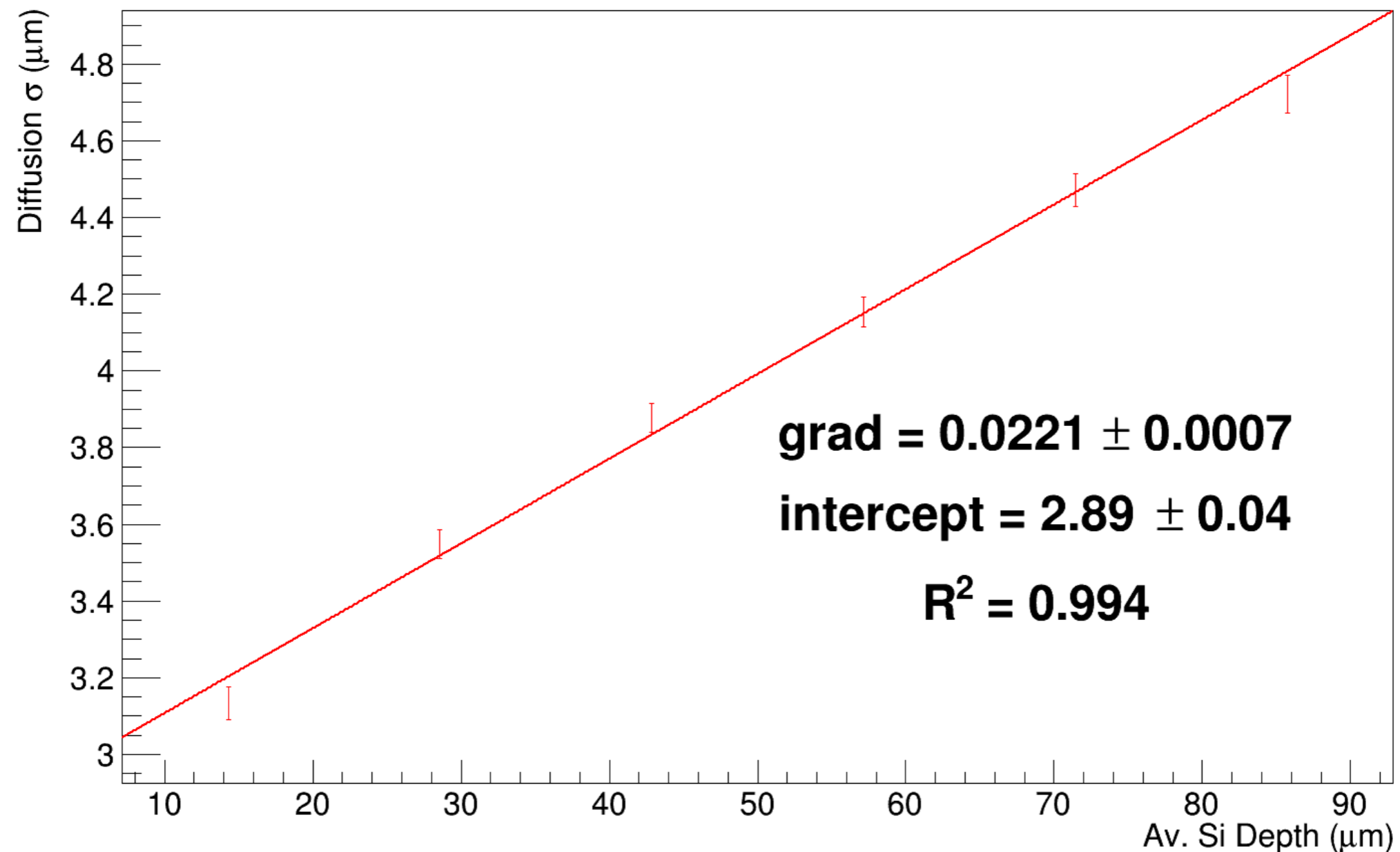
Diffusion Measurement

- ▶ Plot the widths of these Gaussians w.r.t. track segment number
- ▶ Fit a straight line to determine the orientation (directionality) of the track
- ▶ If gradient is positive – leave as is
- ▶ If gradient is negative – reverse order
 - ▶ tracks are now aligned w.r.t sensor surface



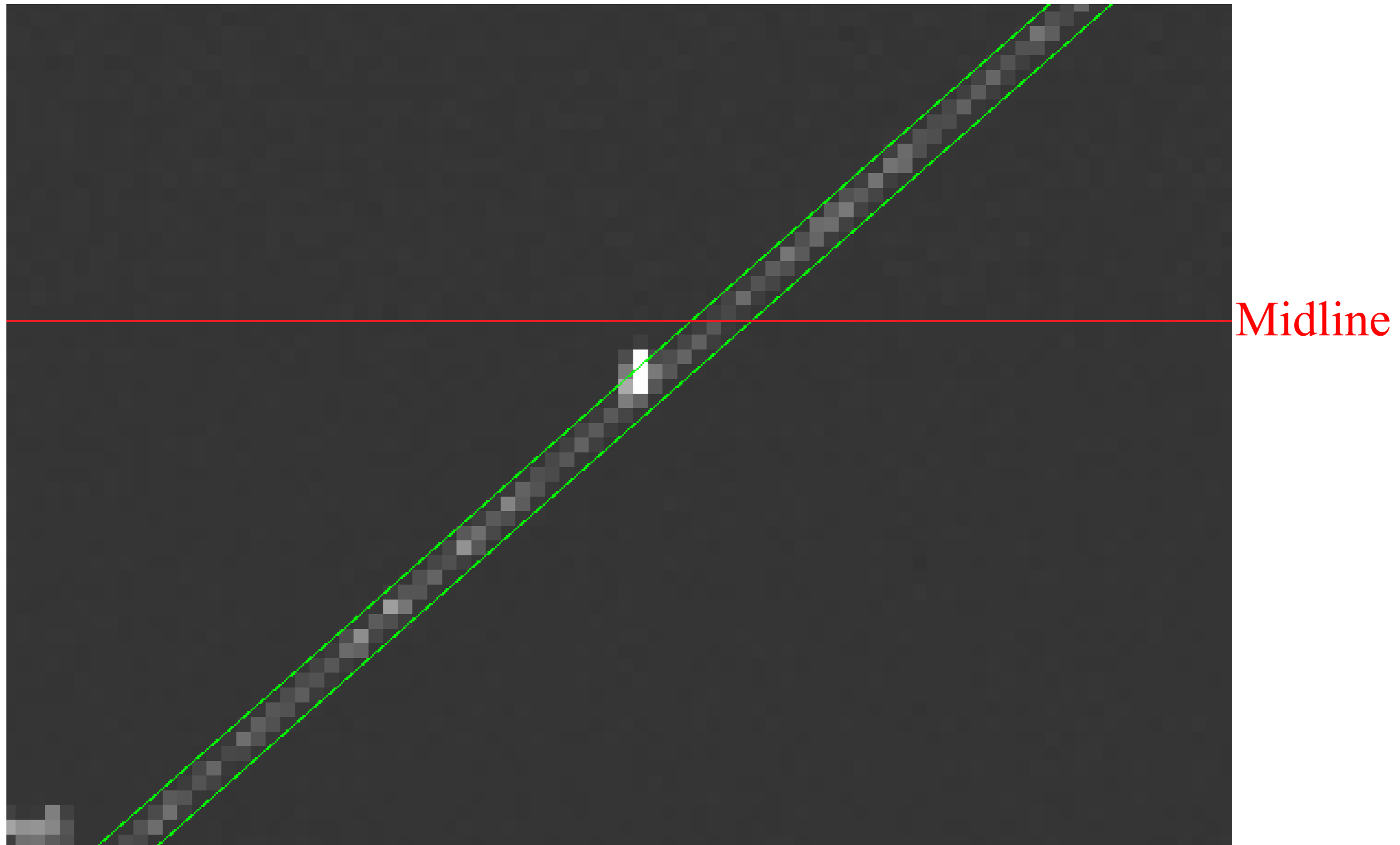
Diffusion Measurement

- Average all tracks segment PSFs together
- Plot these average PSFs as a function of (averaged) depth in silicon
 - → Get diffusion as a function of depth in silicon



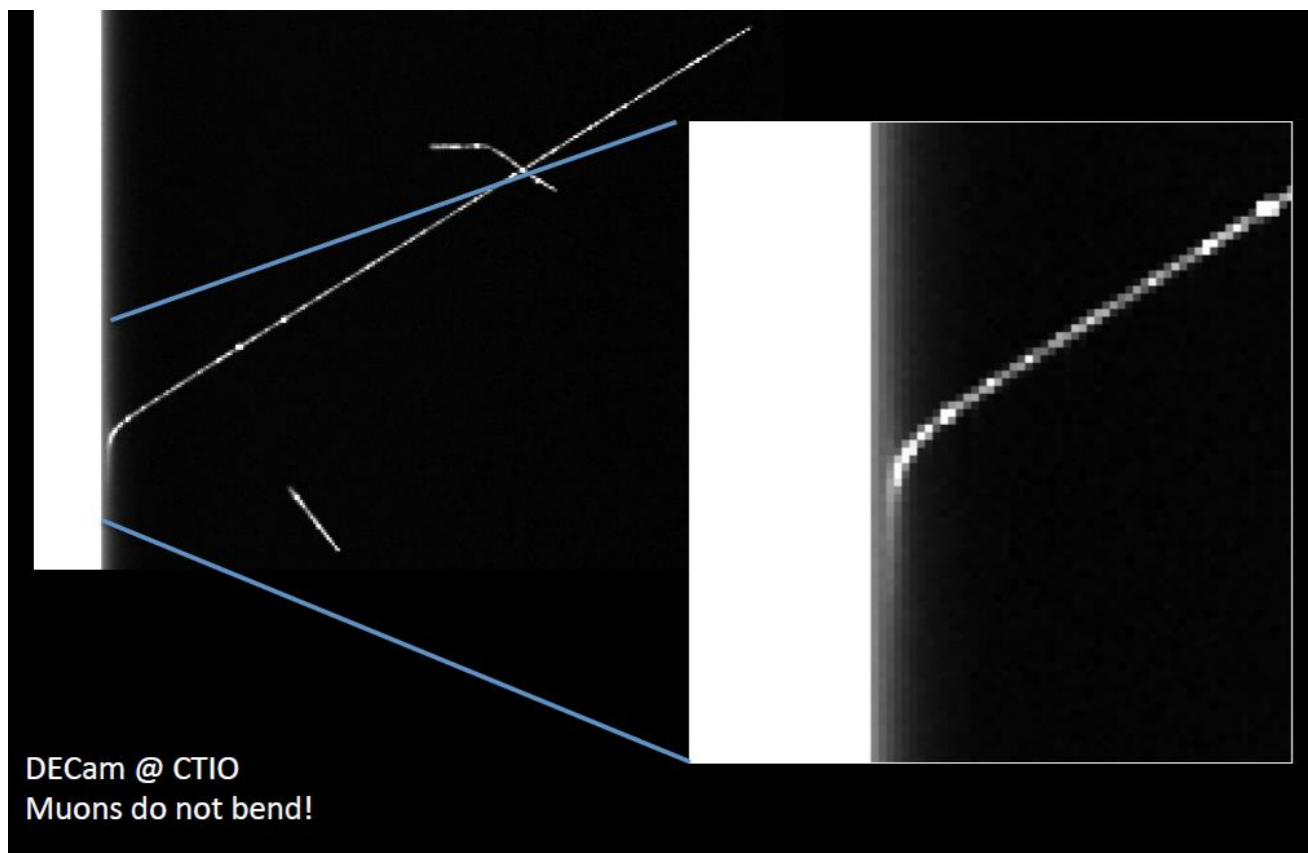
Midline bending

- Midline (not) bending:

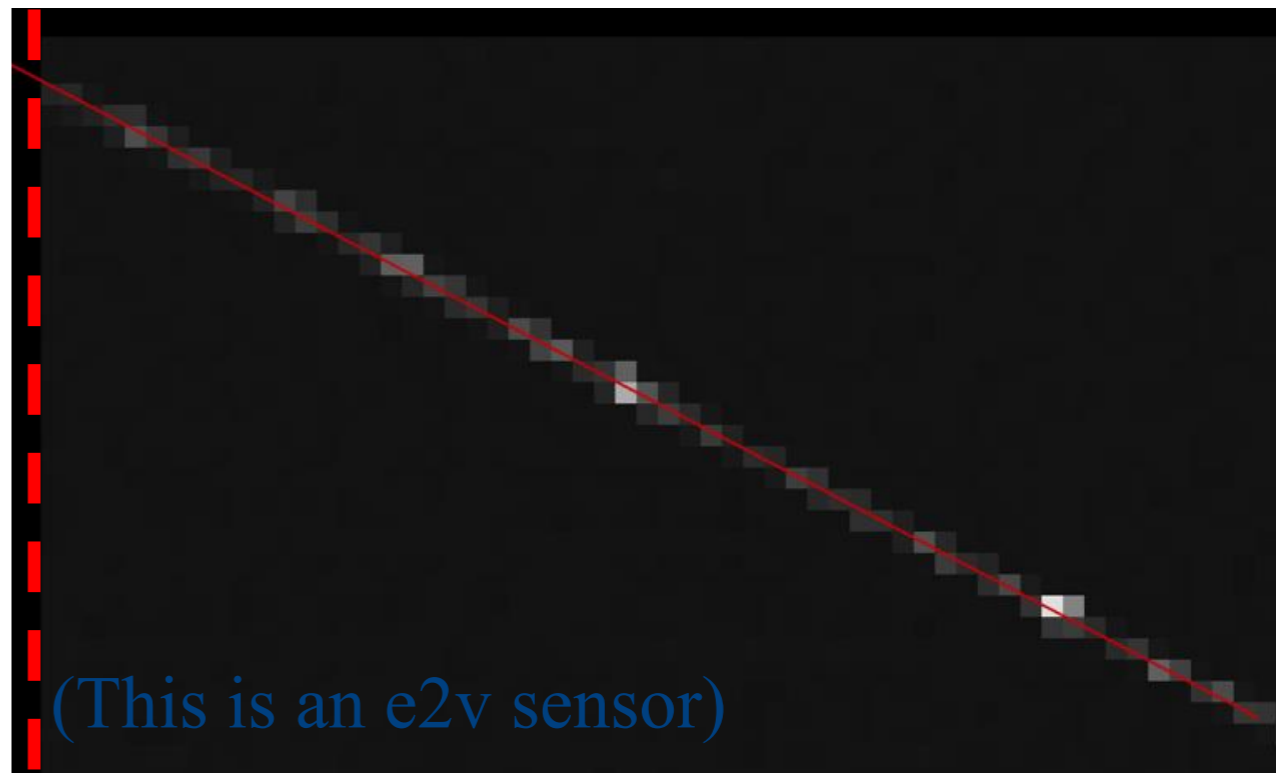


Edge Bending

DECam muon track



LSST muon track



Sensor edge

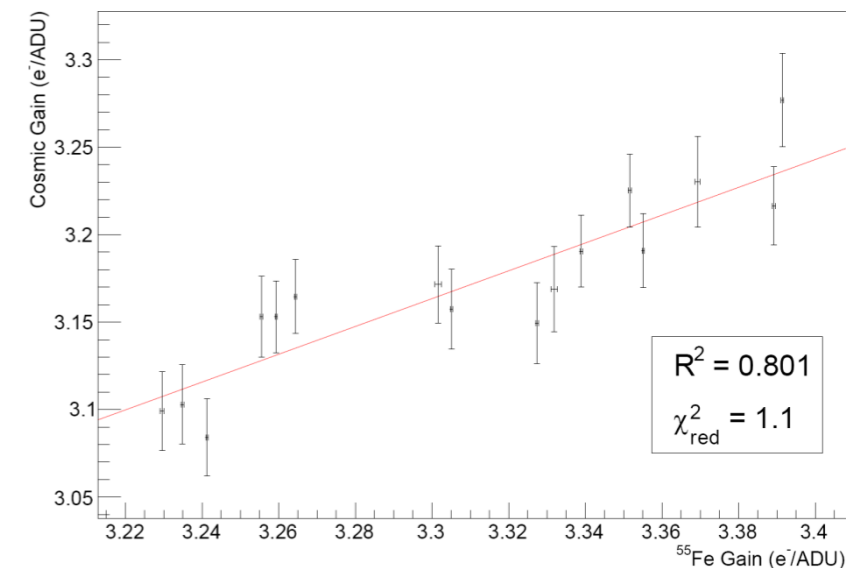
DECam edges both glow
and bend ☹

LSST edges neither glow
nor bend! ☺

Sensor characterisation

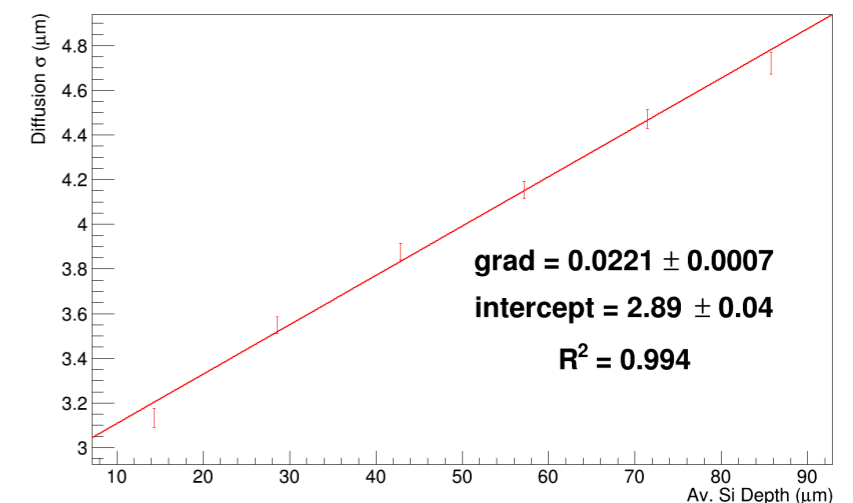
- ▶ Gain measurement with cosmics →

- ▶ Comparison with ^{55}Fe



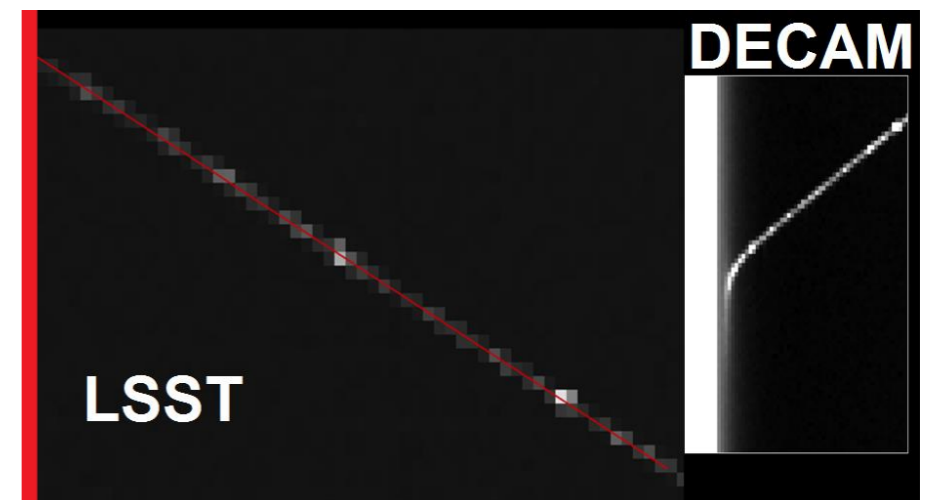
- ▶ Diffusion measurement →

- ▶ PSF found to increase with distance to collection electrode
- ▶ Quantitative analysis to follow



- ▶ Cursory look at edge effects →

- ▶ LSST sensors seem to distort much less than DECcam



Sensor characterisation

► Many thanks to:

- Paul O'Connor & Ivan Kotov for the data
- Robert Lupton & Paul Price for their help with DMStack

► Any questions?

